

Towards non-hydrostatic spectral models

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NCEP EMC Seminar

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NCEP EMC Seminar

Motivations

- Spectral models are used at leading NWP centres.
- Current semi-Lagrangian cores do not make the most out of spectral methods.
- Non-hydrostatic effects need to be evaluated with the non-hydrostatic version rather than a different model.



Three keys

- Accurate associated Legendre functions at high order and degree
- Accurate interpolation in semi-Lagrangian advection
- Stable and simple non-hydrostatic formulation

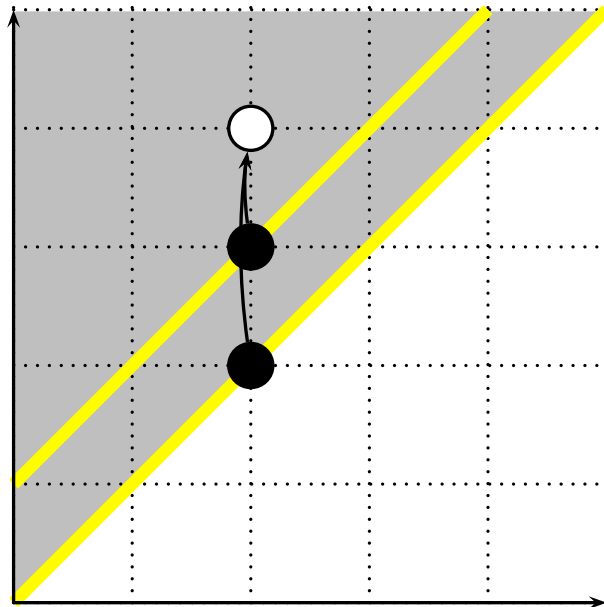


Associated Legendre functions

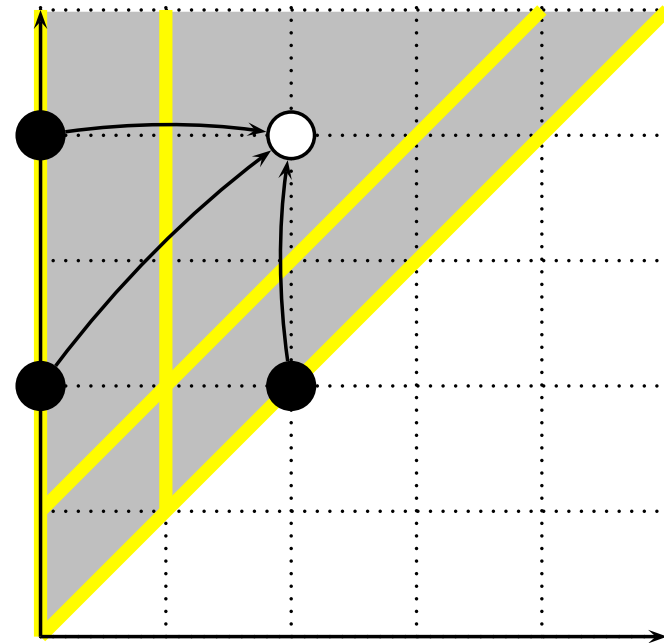
Enomoto et al. (2004; 2008);
Enomoto and Miyamoto, MSJ 2012 Spring



Conventional



Swarztrauber (1993)



Conventional three-term recurrence

$$\tilde{P}_n^m = (a_n^m \cos \theta) \tilde{P}_{n-1}^m - b_n^m \tilde{P}_{n-2}^m \quad (1)$$

$$\tilde{P}_m^m = (d_n^m \sin \theta) \tilde{P}_{m-1}^{m-1} \quad (2)$$

$$\tilde{P}_{m+1}^m = (a_{m+1}^m \cos \theta) \tilde{P}_m^m \quad (3)$$

where θ is the colatitude and

$$a_n^m = \sqrt{\frac{4n^2 - 1}{n^2 - m^2}}, b_n^m = \frac{a_n^m}{a_{n-1}^{m-1}}, d_n^m = \sqrt{\frac{2m + 1}{2m}} \quad (4)$$



The alternative four-term recurrence

involves $(m-2, n)$, $(m-2, n-2)$, and $(m-2, n-2)$ to calculate the value at (m, n) :

$$\begin{aligned}
 \tilde{P}_n^m &= \sqrt{\frac{(2n+1)(n+m-2)(n+m-3)}{(2n-3)(n+m-1)(n+m)}} \tilde{P}_{n-2}^{m-2} \\
 &\quad - \sqrt{\frac{(n-m+1)(n-m+2)}{(n+m-1)(n+m)}} \tilde{P}_n^{m-2} \\
 &\quad + \sqrt{\frac{(2n+1)(n-m)(n-m-1)}{(2n-3)(n+m-1)(n+m)}} \tilde{P}_{n-2}^m
 \end{aligned} \tag{5}$$



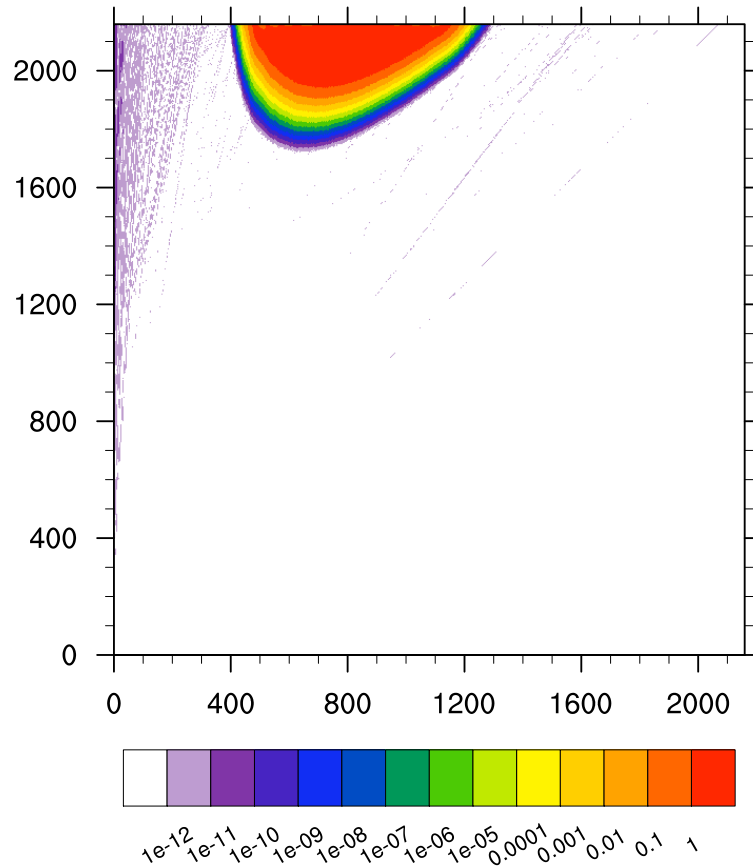
Why does the three-term recurrence fail?

- A floating-point number has can only represent a certain small number e.g. $x_{\min} = 2.23 \times 10^{-308}$ for double precision.
- $\tilde{P}_m^m \propto \sin^m \theta$ is very small near the poles at high orders.
- $|\tilde{P}_n^m|$ grows with degree n .

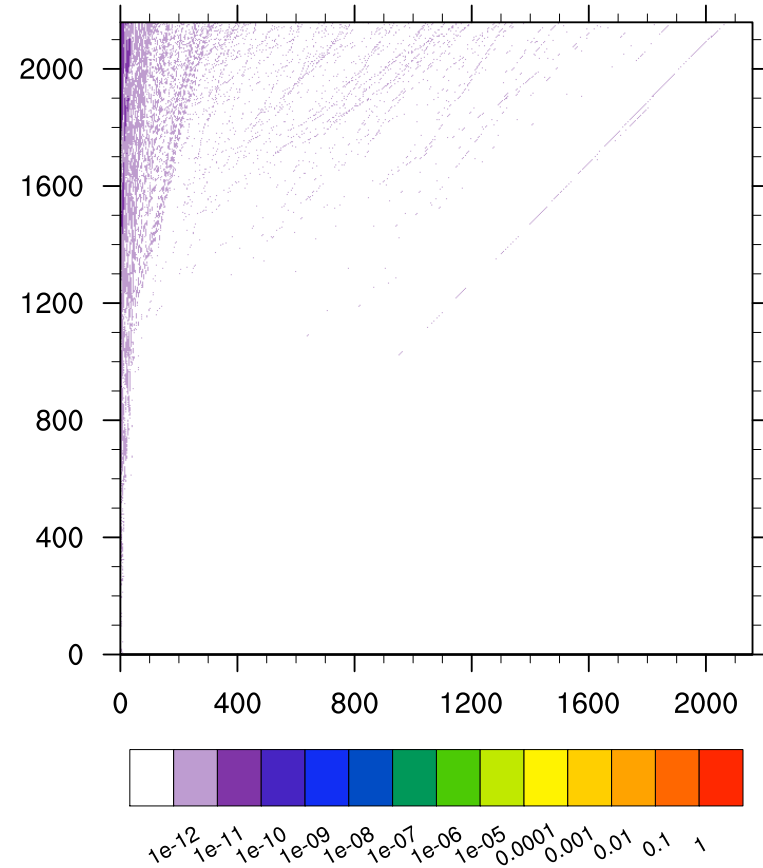


Legendre synthesis–analysis test at T2159

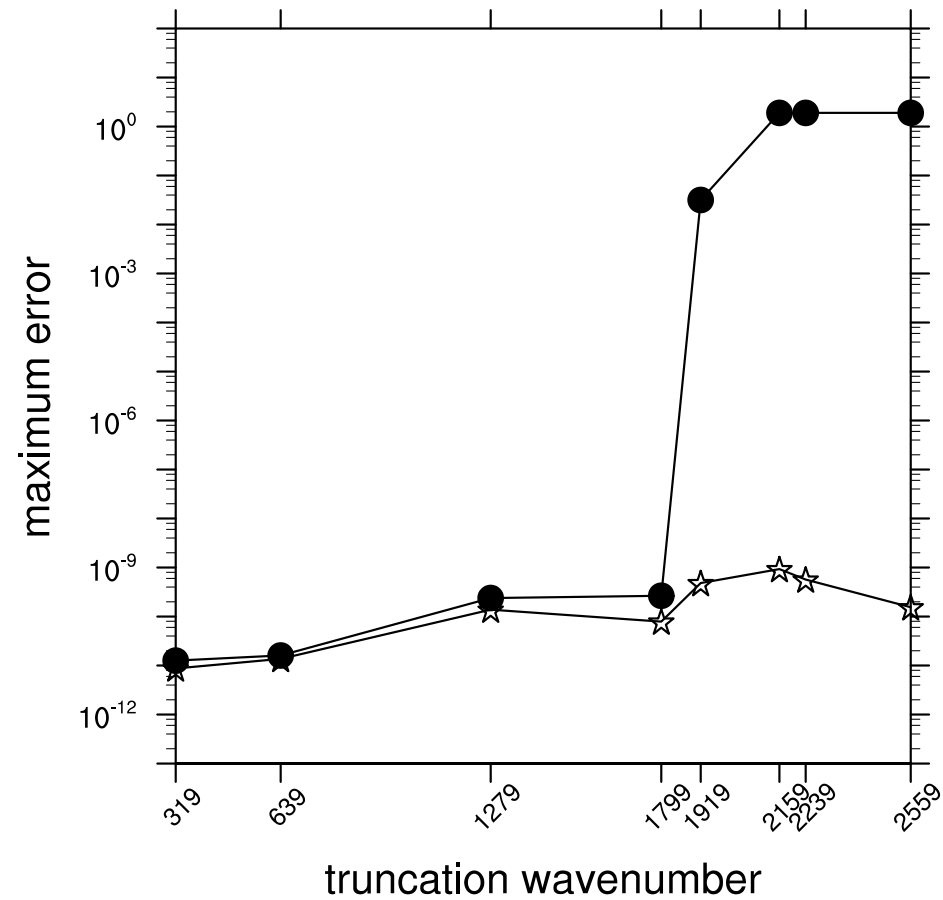
Conventional



Swarztrauber (1993)



Legendre synthesis–analysis test



Extended-range arithmetic

- Smith et al. (1981) avoided overflow and underflow in the evaluation of associated Legendre functions with extended-range arithmetic at the 2x computational cost.
- Fukushima (2011) proposed an accelerated version that enables to evaluate extremely high degree as 2^{32} at only 10% increase in computational time.

Express X with a pair of a floating-point number x and an integer i .

$$X = xB^i \quad (6)$$

where B is a large power of 2.



Validation of associated Legendre functions calculated from the four-term recurrence

- The check sum

$$\int_{-1}^1 [\tilde{P}_n^m]^2 dx = 1 \quad (7)$$

shows that the Swarztrauber (1993)'s recurrence enables accurate Legendre transforms at high order and degree.

- Associated Legendre functions evaluated by the method of Fukushima (2011) is regarded as the truth.
- The small values from the four-term recurrence is found to be inaccurate.



Summary for associated Legendre functions

- The four-term recurrence (Swarztrauber 1993) is stable over 2000.
- The conventional three-term recurrence may be used in extended-range arithmetic to avoid overflow and underflow (Smith et al. 1981).
- An optimized evaluation with extended-range arithmetic only require 10% increase in computational time (Fukushima 2011).
- Associated Legendre functions from the four-term recurrence are validated with those from the three-term recurrence using extended-range arithmetic to be inaccurate at small values.



Interpolation

Enomoto 2008;

Lauritzen et al. 2012, in preparation



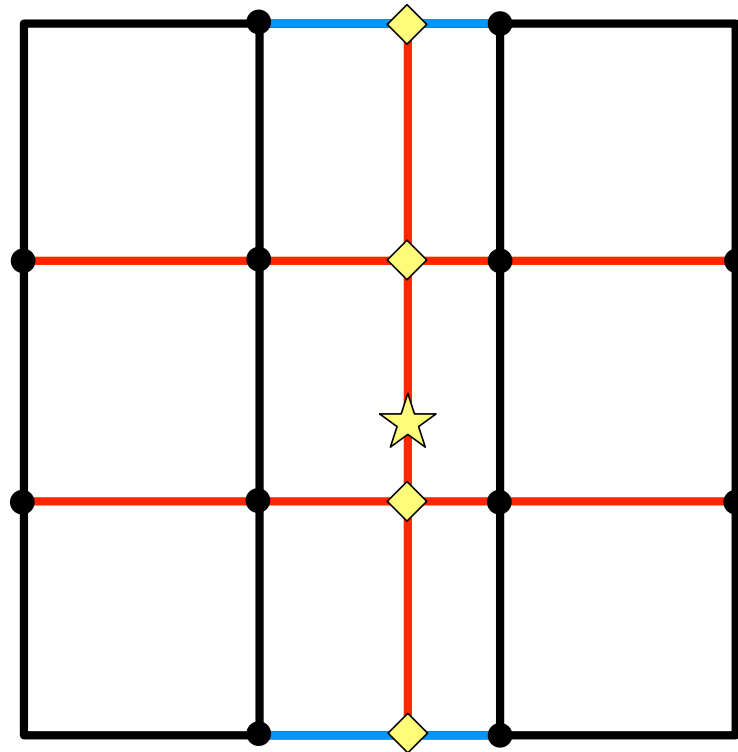
Advection

- Eulerian advection
 - is dispersive
 - requires a short time step.
- Semi-Lagrangian advection
 - allows a longer time step
 - needs interpolation, which determines the accuracy
 - is typically diffusive

→ Semi-Lagrangian advection with an accurate interpolation should be less dispersive and diffusive.



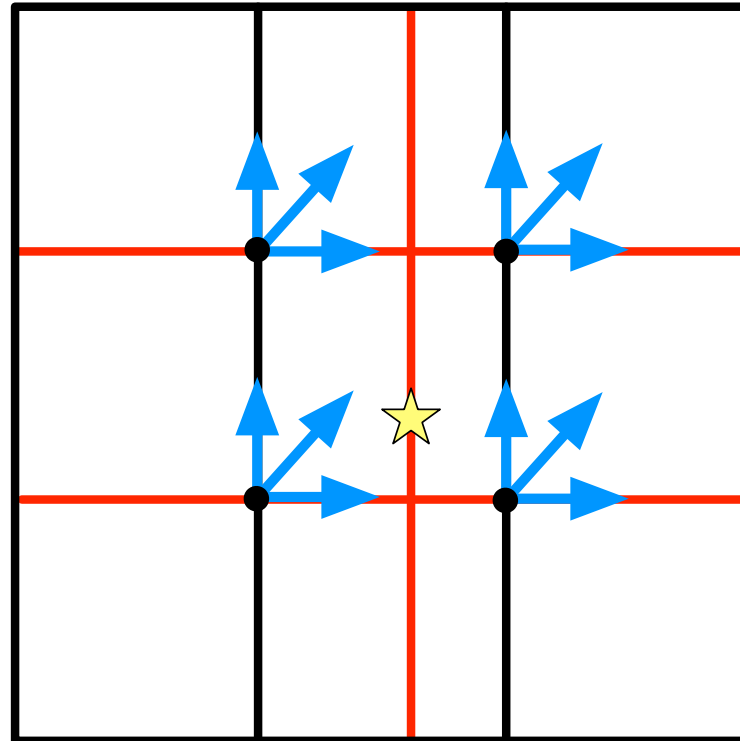
Quasi-cubic interpolation



Ritchie et al. (1995)



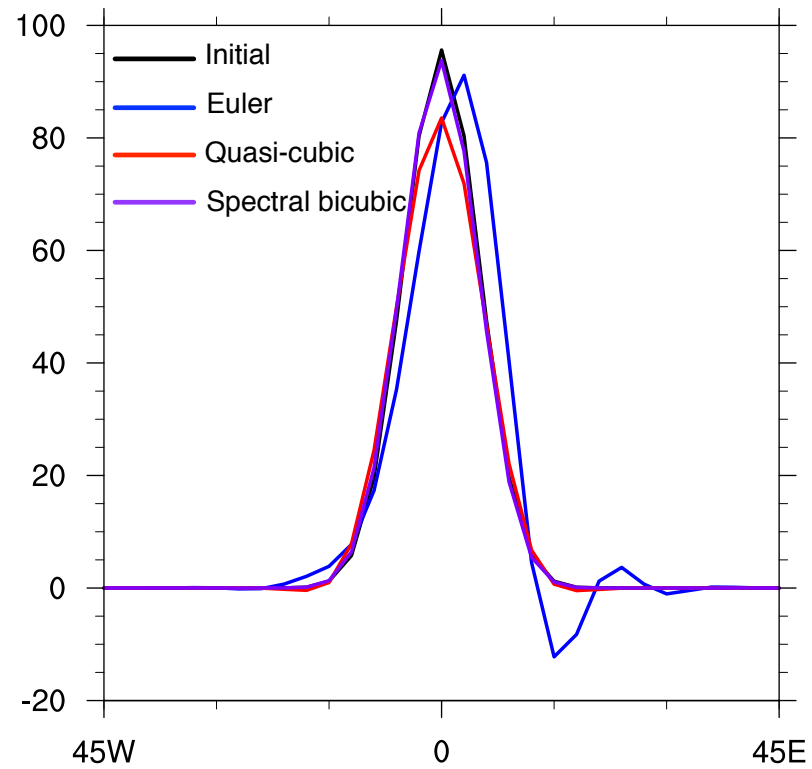
Bicubic interpolation



Derivatives calculated in spectral space (Enomoto 2008)



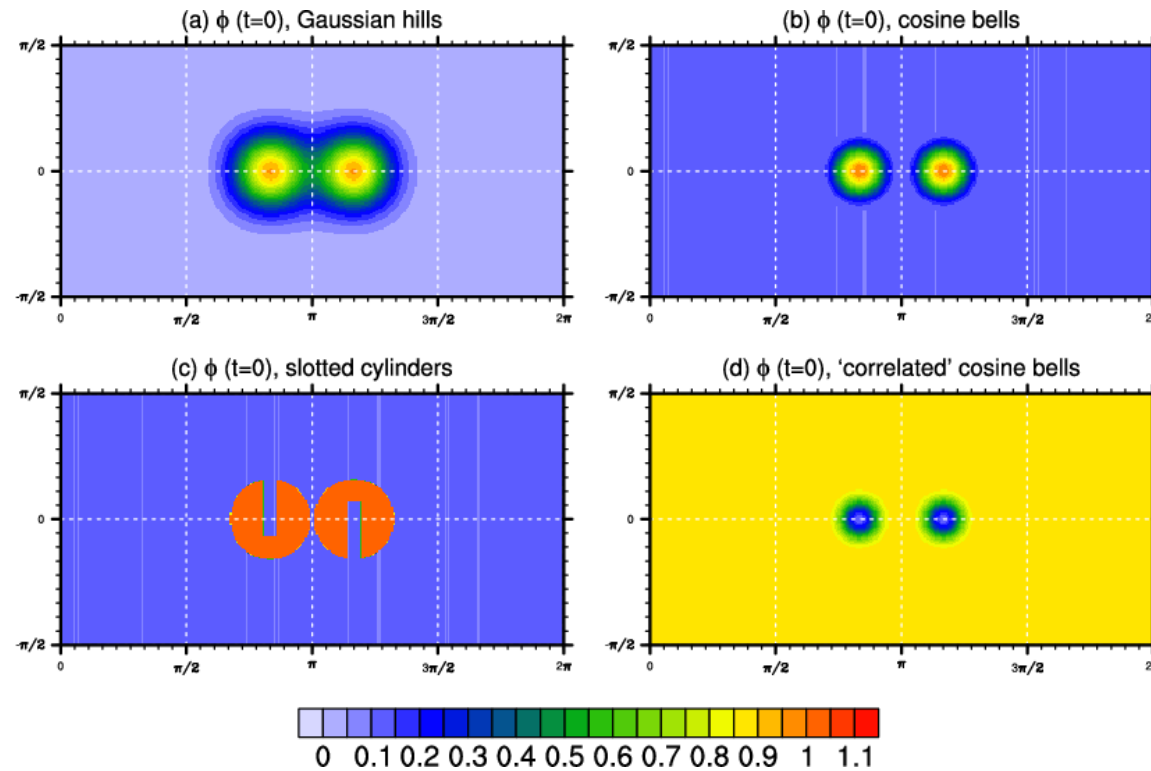
Rotation of a Gaussian hill



Enomoto (2008)



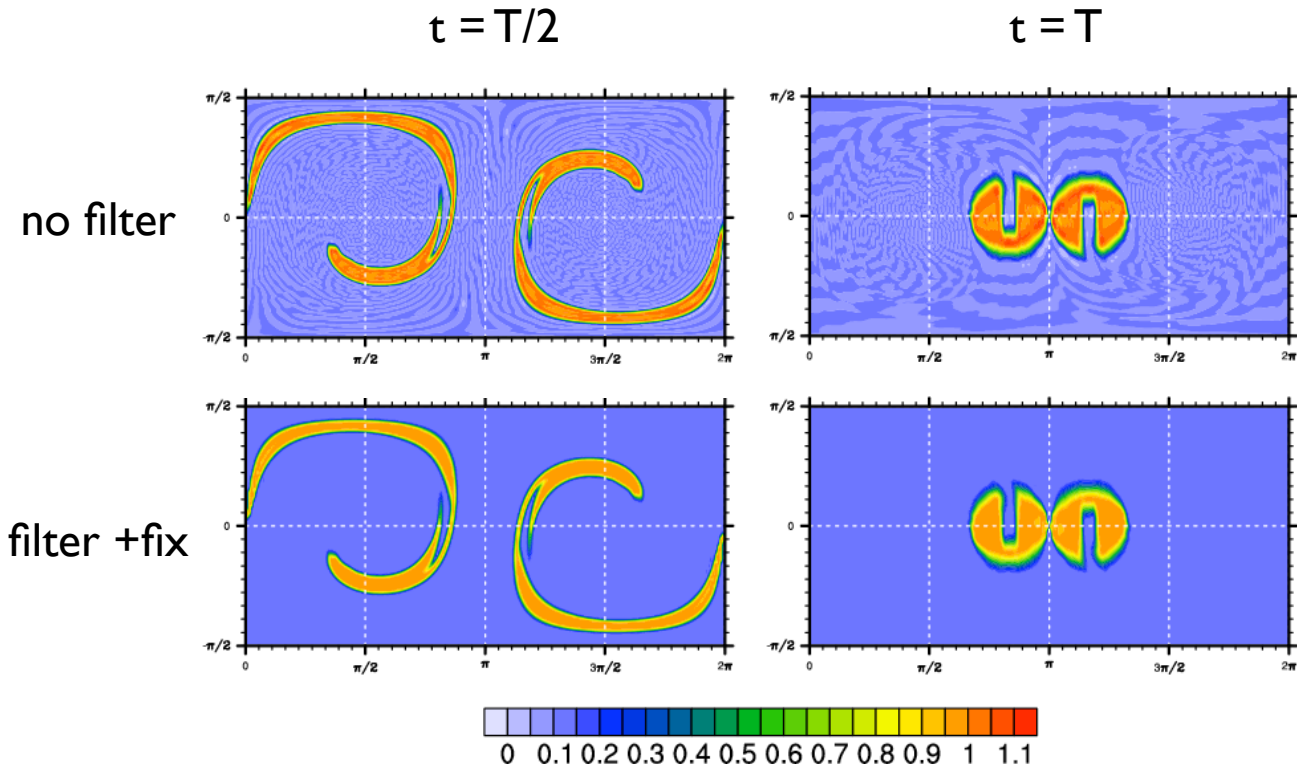
Standard test suite



Lauritzen and Skamarock (2010)



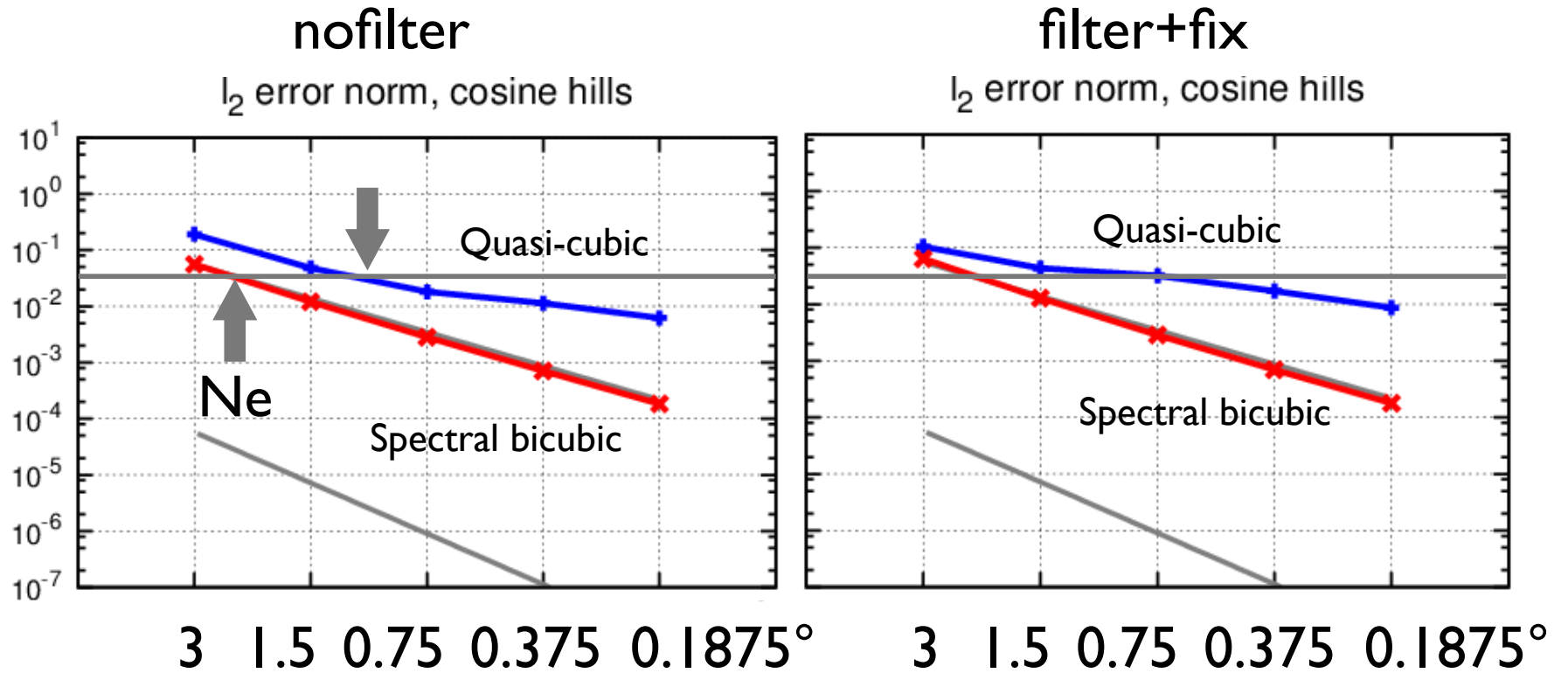
Slotted cylinders



Spectral bicubic T119 CFL 5.2



Convergence



T119 CFL 5.2 l_2 error norm cos hills



'Minimal' resolution compared

model	reference	filter	Ne
Quasi-cubic	Ritchie et al. 1995	no	0.919
		yes	0.864
Spectral bicubic	Enomoto 2008	no	2.41
		yes	2.29

obtained from convergence of l_2 error norm cos hills



Summary for interpolation

Spectral bicubic interpolation

- is easy to implement in existing spectral models.
- is accurate for smooth and non-smooth tracers.
- generates ripples but they can be removed with a short-wave filter of Sun et al. (1996).



Non-hydrostatic formulation

Enomoto and Juang, MSJ 2011 Autumn



σ -co-ordinates in hydrostatic pressure \bar{p} (Laprise 1992)

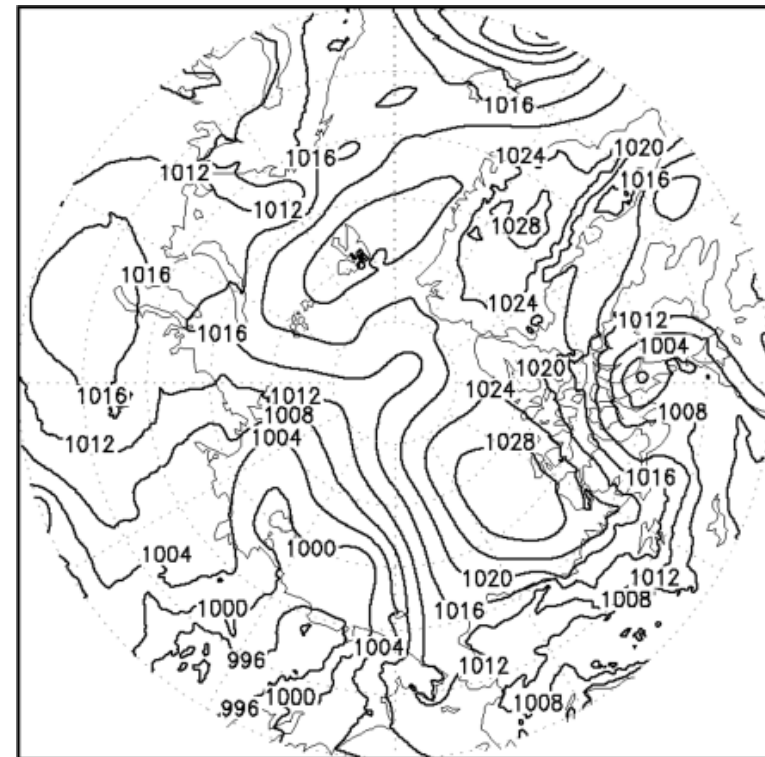
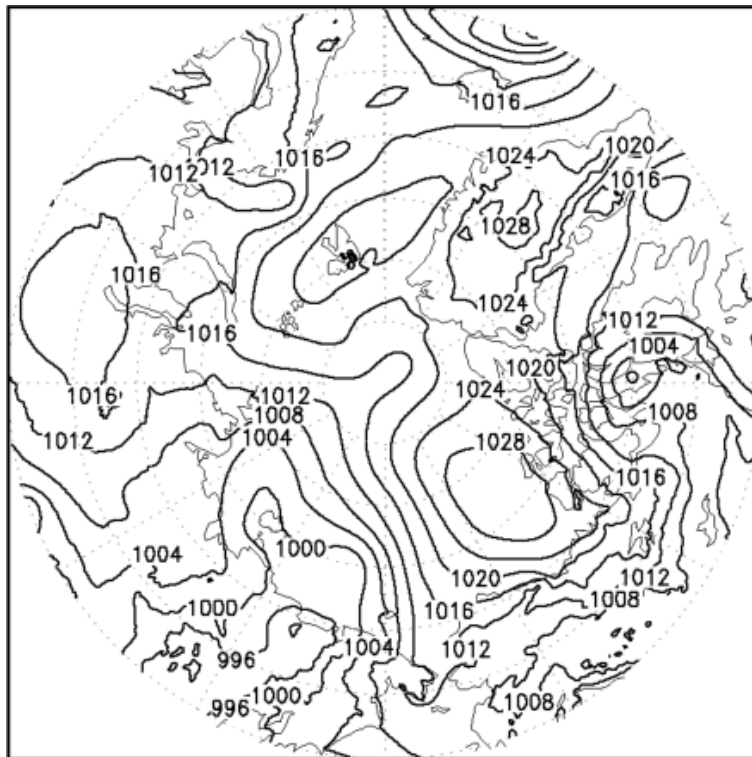
- Additional prognostic variables: w and p
- Monotone in the vertical unlike full pressure p (Miller 1974)
- No additional complex metric terms unlike terrain following height co-ordinates
- Easy to conserve total energy
- Adopted by Météo France ALADIN (Bubnová 1995)/Arpège (Yessard 2008), ECMWF NH-IFS (Wedi et al. 2009) and JMA GSM (Yoshimura, JMSJ spring 2011 B401)



Non-hydrostatic double-fourier version of JMA GSM (Yoshimura)

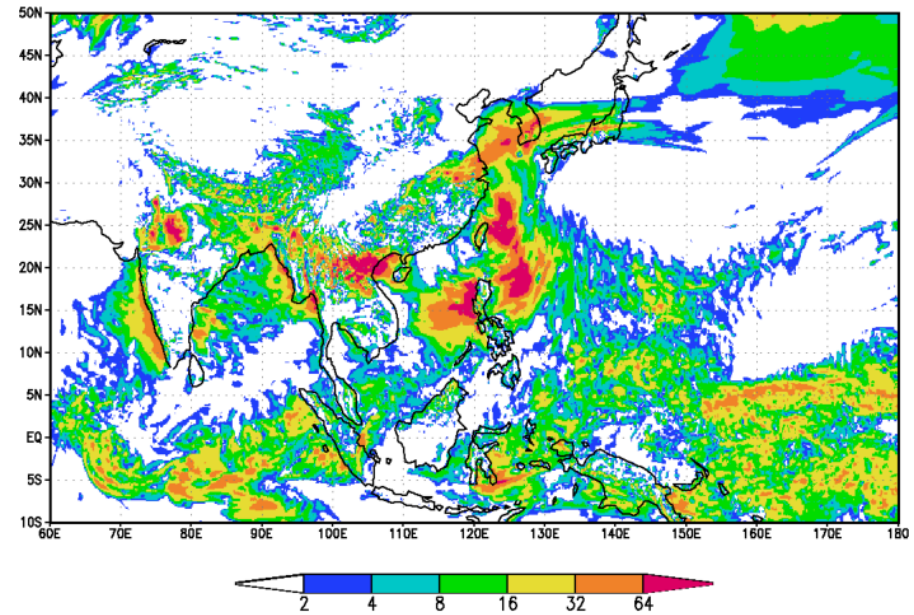
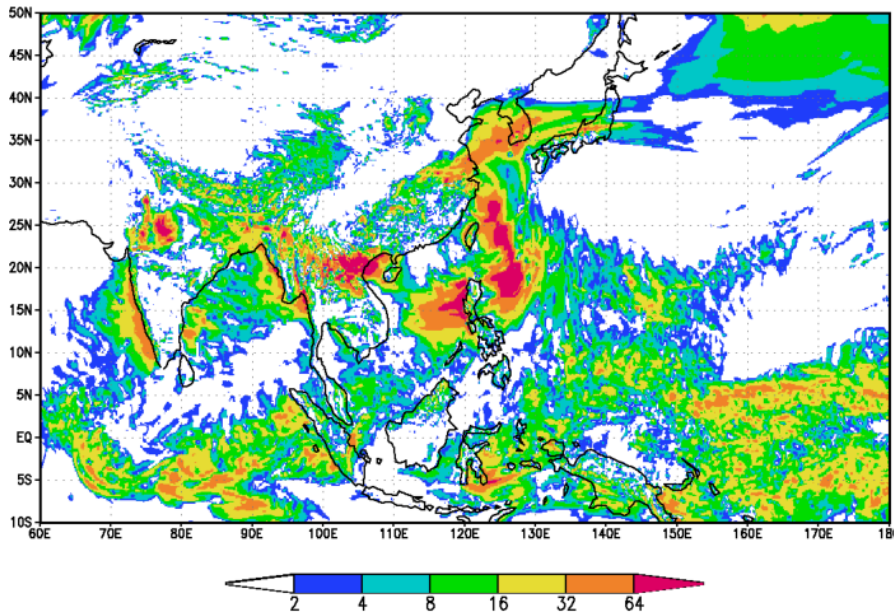
TL1279L60, $\Delta t=10$ min, FT = 2 d, slp (hPa)
Spherical harmonics

Double Fourier Series

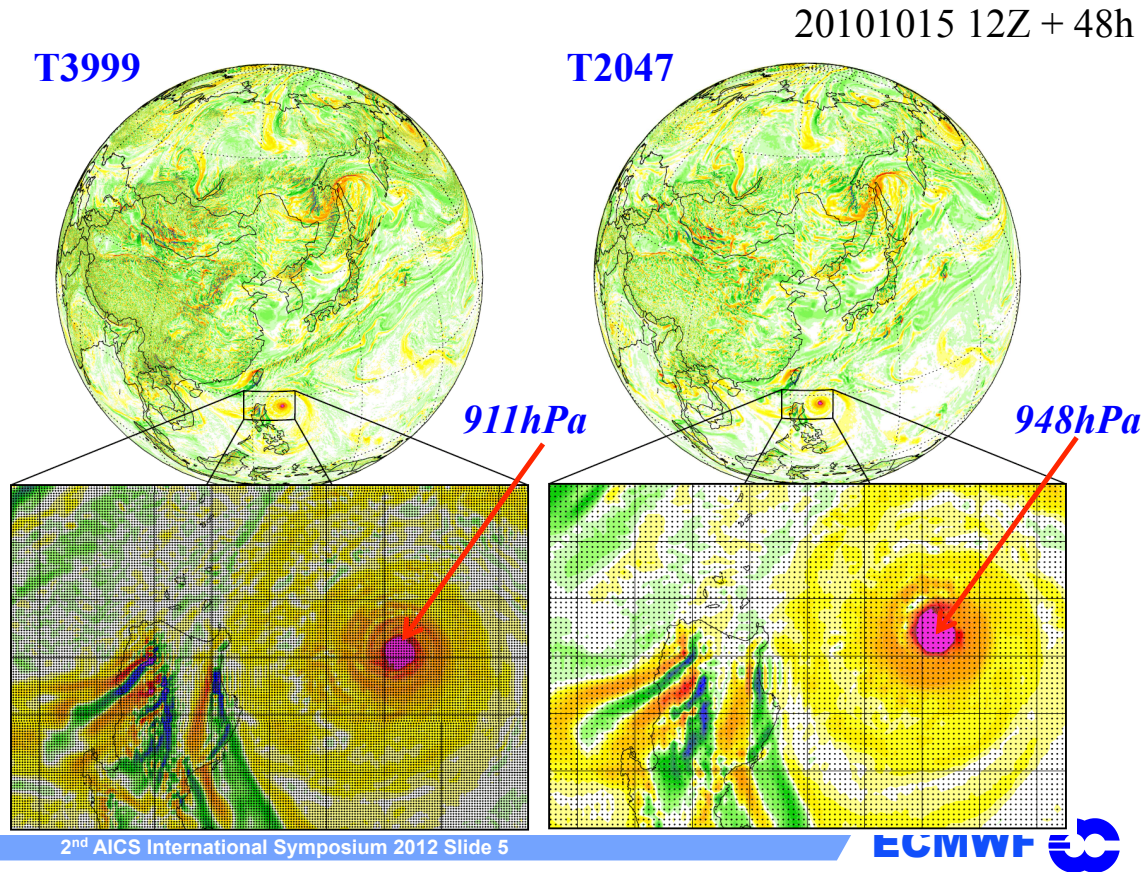


Non-hydrostatic double-fourier version of JMA GSM (Yoshimura)

TL1279L60, $\Delta t=10$ min, FT = 2 d, precipitation (mm/d)
Spherical harmonics Double Fourier Series



Non-hydrostatic IFS at ECMWF (Wedi)



You already have
a gem at NCEP.



NCEP MSM (Juang 1992; 2000)

- is a non-hydrostatic version of NCEP RSM (Juang and Kanamitsu 1994).
- uses horizontal discretization by double Fourier series.
- uses the perturbation method.
- shares the same physics packages with NCEP GSM.
- uses the Euler equations transformed from z - to σ - co-ordinates.



Pressure gradient terms

Laprise (1992)

$$\frac{1}{\rho} \nabla_z p = RT \nabla_\sigma \ln p + \frac{p \partial \ln p}{\bar{p} \partial \ln \sigma} \nabla_\sigma \phi, \quad \frac{\partial \phi}{\partial \sigma} = -\frac{RT \bar{p}_s}{p} \quad (8)$$

Juang (1992)

$$\frac{1}{\rho} \nabla_z p = RT \nabla_\sigma \ln p + \frac{T \partial \ln p}{\bar{T} \partial \ln \sigma} \nabla_\sigma \bar{\phi}, \quad \frac{\partial \bar{\phi}}{\partial \sigma} = -\frac{RT}{\sigma} \quad (9)$$



The hydrostatic state of Laprise (1992)

Define

$$\frac{\partial \phi}{\partial p} = -\frac{1}{\rho}, \quad (10)$$

means

$$\bar{\rho} = \rho. \quad (11)$$

Since $\bar{\rho} = \bar{p}/R\bar{T}$, $\rho = p/RT$,

$$\frac{T}{\bar{T}} = \frac{p}{\bar{p}} \quad (12)$$



Hydrostatic variables

Choice of the hydrostatic temperature \bar{T} and surface pressure \bar{p}_s :

1. Time independent (Juang 1994; Gallus and Rančić 1996)
2. Both \bar{T} and \bar{p}_s are determined externally (Juang 1992).
3. Impose \bar{T} externally but predict \bar{p}_s internally (Juang 2000, default of MSM).
4. Predict both \bar{T} and \bar{p}_s internally (Juang 2000).



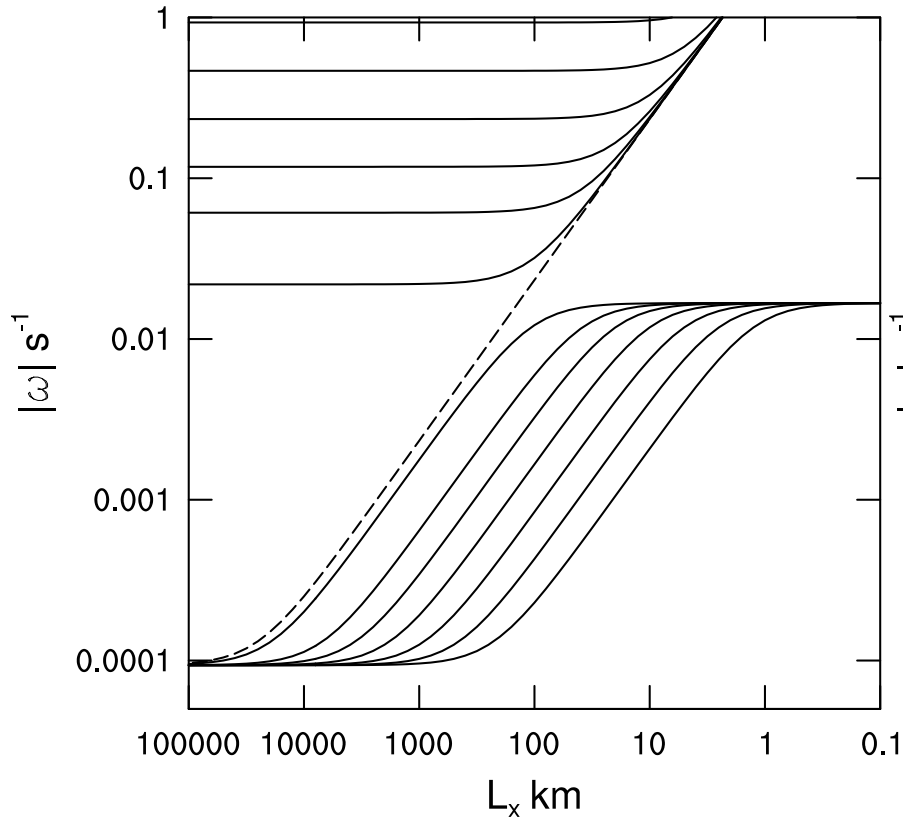
The unified system by Arakawa and Konor (2009)

- unifies anelastic and primitive (quasi-hydrostatic) systems.
- uses the hydrostatic density $\bar{\rho}$ in the continuity equation.
 - by ignoring non-hydrostatic pressure tendency $\partial(p - \bar{p})/\partial t$:
 $p - \bar{p}$ is obtained by solving an Helmholtz equation.
- removes sound waves without distortion of planetary waves.

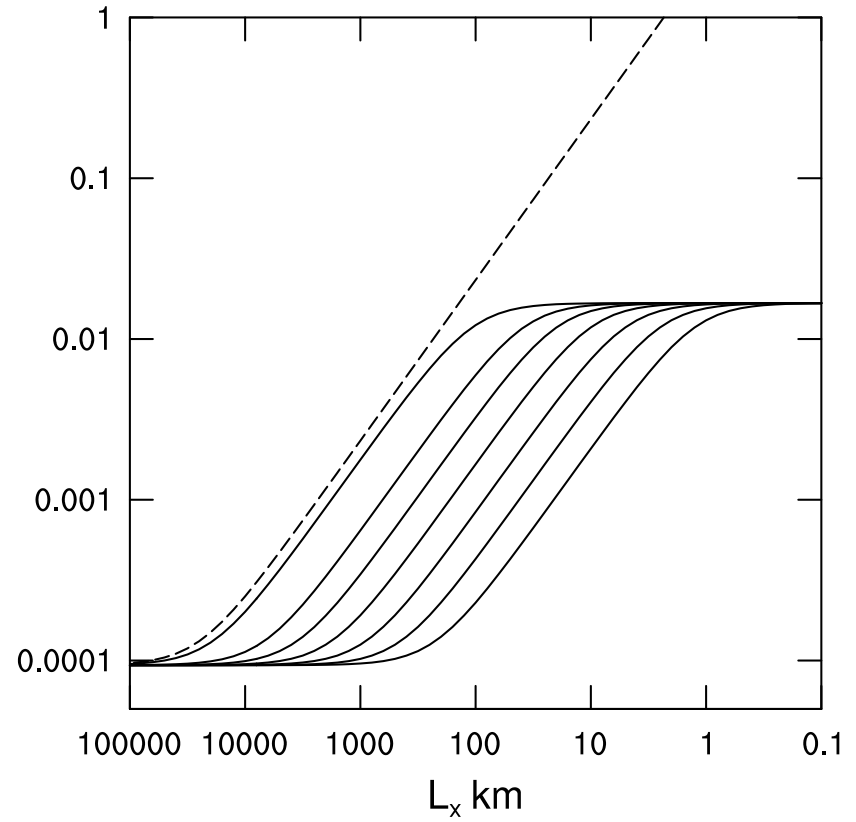


Sound and gravity waves

Fully-Compressible

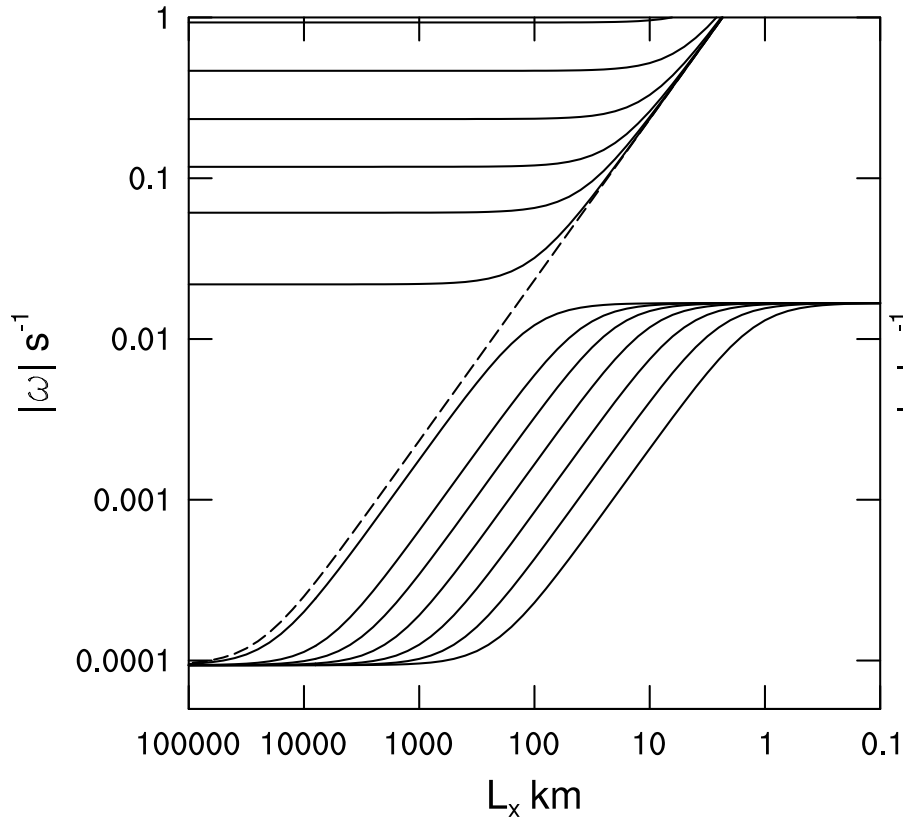


Unified

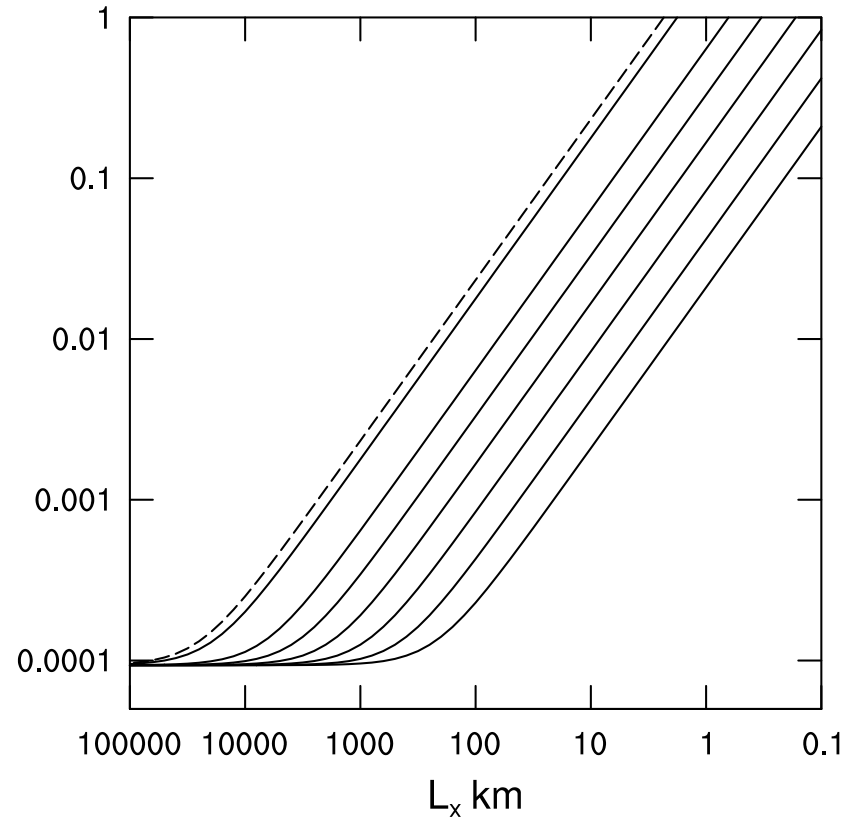


Sound and gravity waves

Fully-Compressible

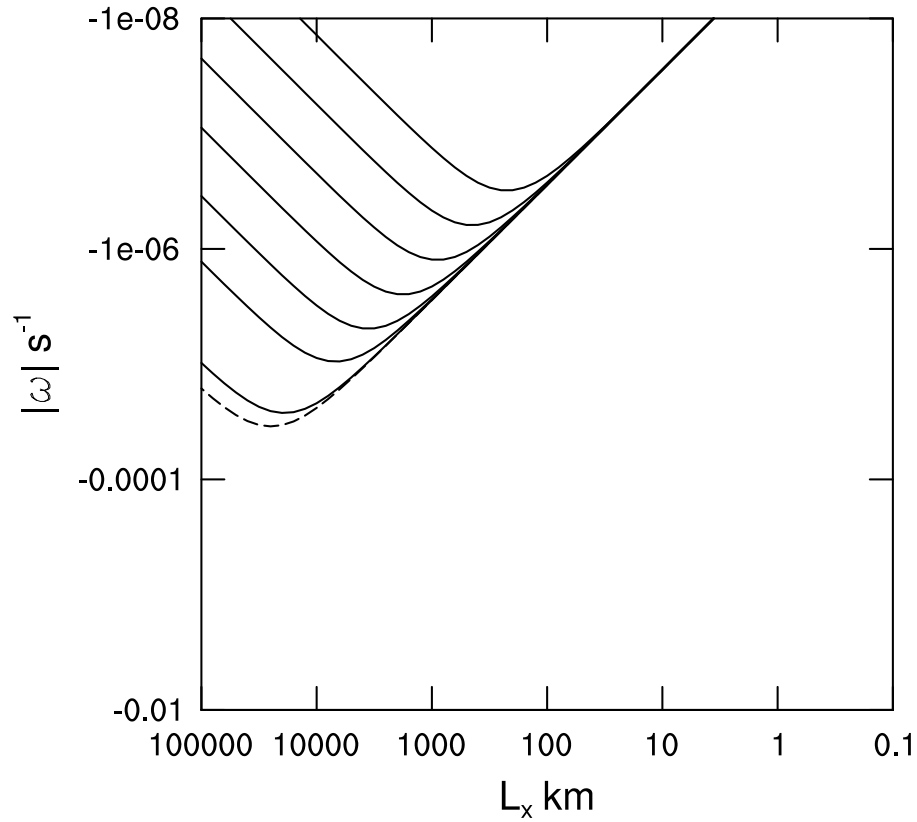


Quasi-Hydrostatic

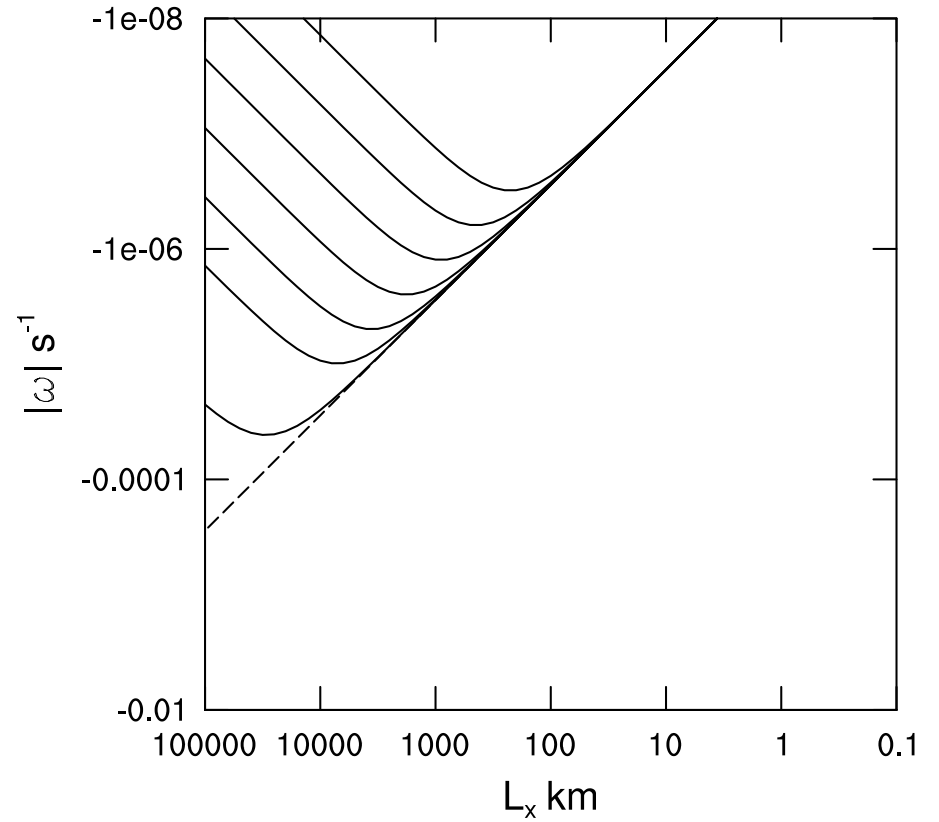


Rossby waves

Fully-Compressible, Unified, Quasi-Hydrostatic



Pseudo-Incompressible



The hydrostatic state of Arakawa and Konor (2009)

Using the hydrostatic Exner function $\bar{\pi} = (\bar{p}/p_{\text{ref}})^\kappa$, define

$$\frac{\partial \bar{\pi}}{\partial z} \equiv -\frac{g}{c_p \bar{\theta}}, \quad (13)$$

which means

$$\bar{\theta} = \theta. \quad (14)$$

Since $\theta = T/\pi$ and $\bar{\theta} = \bar{T}/\bar{\pi}$, hydrostatic temperature \bar{T} may be written as

$$\bar{T} = T \left(\frac{\bar{p}}{p} \right)^\kappa. \quad (15)$$

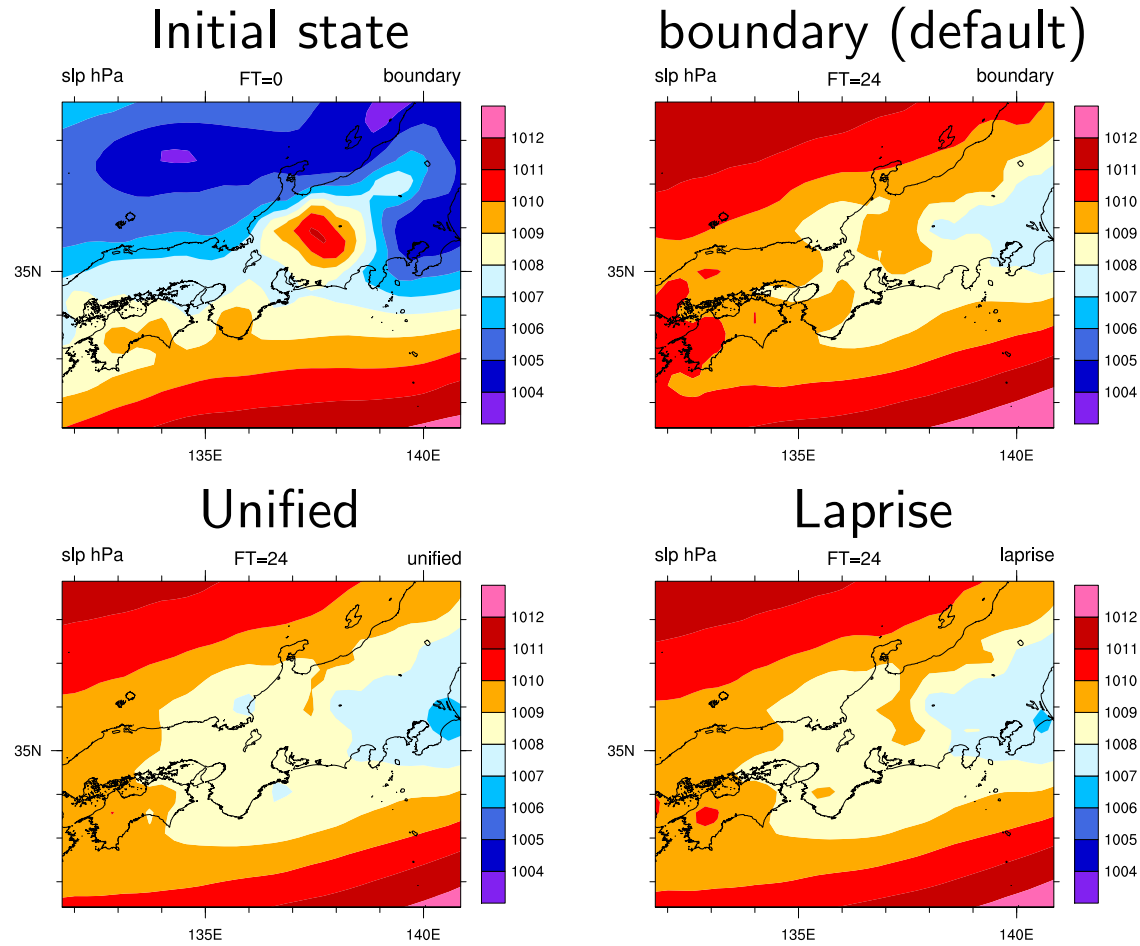


Forecast experiment using NCEP MSM

- Initial time: 0 UTC 11 August 2011, 24-hour forecast
- Horizontal resolution: $\Delta h = 26$ km, vertical levels: 42, time step: $\Delta t = 60$ s
- Initial and boundary conditions: NCEP GFS
- Projection: polar stereo
- Domain: (132–141E, 31–39N) centred at (136E, 35N)
- Problem size: $32 \times 32 \times 42$



slp hPa FT=24 INIT=2011081800



Summary for non-hydrostatic formulation

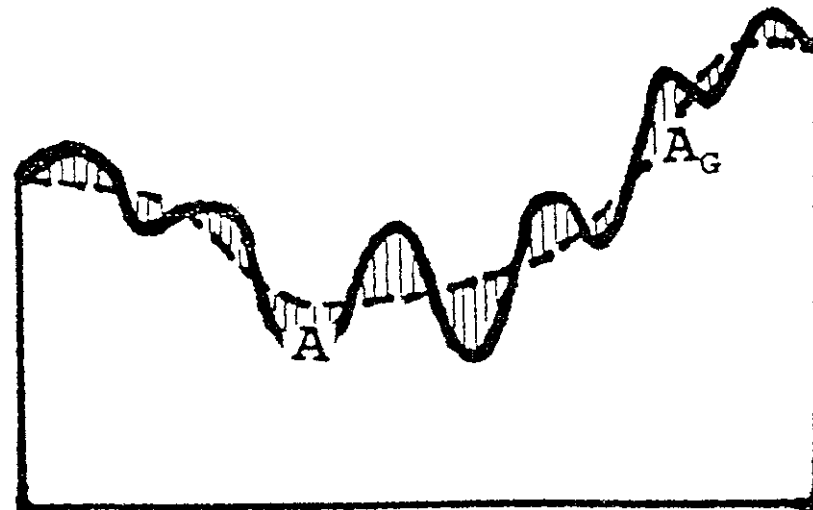
- The governing equation of MSM is the Euler equations in σ -coordinates transformed from those in z -coordinates (Juang 1992).
- Sound wave can be removed without distorting planetary wave by the use of the hydrostatic density $\bar{\rho}$ in the continuity equation (Arakawa and Konor 2009).
- The hydrostatic temperature state \bar{T} may be defined by the hydrostatic assumption of Arakawa and Konor (2009) or Laprise (1992).
- Preliminary forecast experiments shows that both hydrostatic assumption are stable.





The perturbation method

predicts deviations from from the global model.



PERTURBATION METHOD

$$\frac{\partial A'}{\partial t} = \frac{\partial A}{\partial t} - \frac{\partial A_b}{\partial t} \quad (16)$$



The governing equations

$$\begin{aligned}
 \frac{\partial \mathbf{v}'}{\partial t} = & -m^2 \mathbf{v} \cdot \nabla \mathbf{v} - \dot{\sigma} \frac{\partial \mathbf{v}}{\partial \sigma} - E \nabla m^2 + f \mathbf{k} \times \mathbf{v} \\
 & - R(\bar{T} + T') \nabla (\bar{Q}_s + Q') \\
 & - \left(1 + \frac{T'}{\bar{T}}\right) \left(1 + \frac{\partial Q'}{\partial \ln \sigma}\right) \nabla \bar{\phi} + \mathbf{F} - \frac{\partial \mathbf{v}_b}{\partial t}
 \end{aligned} \tag{17}$$

$$\begin{aligned}
 \frac{\partial w'}{\partial t} = & -m^2 \mathbf{u} \cdot \nabla w - \dot{\sigma} \frac{\partial w}{\partial \sigma} \\
 & - g \left[1 - \left(1 + \frac{T'}{\bar{T}}\right) \left(1 + \frac{\partial Q'}{\partial \ln \sigma}\right)\right] + F_z - \frac{\partial w_b}{\partial t}
 \end{aligned} \tag{18}$$



$$\frac{\partial \bar{Q}'_s}{\partial t} = -m^2 \int_0^1 \left[u \frac{\partial \bar{Q}_s}{\partial x} + v \frac{\partial \bar{Q}_s}{\partial y} + \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) \right] d\sigma - \frac{\partial Q_{sb}}{\partial t} \quad (19)$$

$$\begin{aligned} \frac{\partial Q'}{\partial t} = & -m^2 u \frac{\partial \bar{Q}_s + Q'}{\partial x} - m^2 v \frac{\partial \bar{Q}_s + Q'}{\partial y} - \dot{\sigma} \frac{\partial Q'}{\partial \sigma} - \frac{\dot{\sigma}}{\sigma} \\ & - \gamma \nabla_3 \cdot \mathbf{v} + \gamma \frac{F_T}{T} - \frac{\partial \bar{Q}_s}{\partial t} \end{aligned} \quad (20)$$

$$\begin{aligned} \frac{\partial T'}{\partial t} = & -m^2 u \frac{\partial \bar{T} + T'}{\partial x} - m^2 v \frac{\partial \bar{T} + T'}{\partial y} - \dot{\sigma} \dot{\sigma}^\kappa \frac{\partial (\bar{T} + T') \dot{\sigma}^{-\kappa}}{\partial \sigma} \\ & - \frac{RT}{c_v} \nabla_3 \cdot \mathbf{v} + F_T - \frac{\partial \bar{T}_b}{\partial t} \end{aligned} \quad (21)$$

$$\frac{\partial q'}{\partial t} = -m^2 u \frac{\partial q}{\partial x} + m^2 v \frac{\partial q}{\partial y} - \dot{\sigma} q \sigma + F_q - \frac{\partial q_b}{\partial t} \quad (22)$$



Vertical velocity and divergence

$$\dot{\sigma} = \frac{\sigma}{RT} \left[gw + \frac{\partial \bar{\phi}}{\partial t} + m^2 \left(u \frac{\partial \bar{\phi}}{\partial x} + v \frac{\partial \bar{\phi}}{\partial y} \right) \right] \quad (23)$$

$$\nabla_3 \cdot \mathbf{v} = m^2 \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) + \frac{\sigma}{RT} \left[m^2 \left(\frac{\partial u}{\partial \sigma} \frac{\partial \bar{\phi}}{\partial x} + \frac{\partial v}{\partial \sigma} \frac{\partial \bar{\phi}}{\partial y} \right) - g \frac{\partial w}{\partial \sigma} \right] \quad (24)$$

$$\bar{\phi} = \bar{\phi}_s + \int_{\sigma}^{\sigma_s=1} \frac{RT}{\sigma} d\sigma \quad (25)$$



$$Q' \equiv \ln p - \ln \bar{p} = Q - \bar{Q}_s - \ln \sigma \quad (26)$$

gives

$$\left(\frac{\bar{p}}{p}\right)^\kappa = \exp(-\kappa Q'). \quad (27)$$

Using this identity

$$\begin{aligned} \nabla \bar{T} &= \left(\frac{\bar{p}}{p}\right)^\kappa (\nabla T - \kappa T \nabla Q') \\ &= \exp(-\kappa Q') (\nabla T - \kappa T \nabla Q'). \end{aligned} \quad (28)$$

$\nabla \bar{T}$ is used in $\dot{\sigma}$ and $\nabla_3 \cdot \mathbf{v}$.

