

1 **Ensemble Transform with 3D Rescaling Initialization Method**

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3
4 Juhui Ma¹, Yuejian Zhu², Dingchen Hou², Xiaqiong Zhou³ and Malaquias Peña³

5 ¹ Key Laboratory of Meteorological Disaster of Ministry of Education,
6 Nanjing University of Information Science and Technology, Nanjing 210044, China

7 ² Environmental Modeling Center/NCEP/NOAA, College Park, MD 20740, USA

8 ³ I.M. Systems Group, Inc. (IMSG) at the Environmental Modeling Center/NCEP/NOAA,
9 College Park, MD 20740, USA

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19 Corresponding author address:

20 Juhui Ma

21 1100 Perimeter Park Dr #104

22 Morrisville, NC 27560

23 Email: juhui.ma@gmail.com

24

25 **Abstract**

26 The Ensemble Transform with Rescaling (ETR) method has been used to produce fast
27 growing components of analysis error in the NCEP Global Ensemble Forecast System (GEFS).
28 The rescaling mask contained in the ETR method constrains the amplitude of perturbations to
29 reflect regional changes of analysis error. However, due to a lack of suitable three-dimensional
30 (3D) analysis error estimation, in the operational GEFS the mask is based on an estimated
31 analysis error at 500hPa and is not flow-dependent but changes monthly. With the availability of
32 an ensemble-based data assimilation system at NCEP, a 3D mask can be computed. This study
33 generates initial perturbations by the Ensemble Transform with 3D Rescaling (ET_3DR) and
34 compares the performance with the ETR. Meanwhile, the ET_3DR is also applied into the
35 Ensemble Kalman Filter (EnKF) method (hereinafter referred to as EnKF_3DR).

36 Results from a set of experiments indicate that the 3D mask affects the amplitude of initial
37 perturbations. Relative to the ETR, the large amplitudes of the ET_3DR initial perturbations at
38 500hPa connect better with baroclinic instability areas over the extra-tropics and deep convection
39 areas over the tropics. Furthermore, the maxima of vertical distribution for the ET_3DR initial
40 perturbations correspond to the subtropical jet region and tropical easterlies jet region. The better
41 distribution of the perturbations is found to produce faster spread growths. Results with
42 EnKF_3DR also show benefits. The variance along orthogonal basis vectors in the EnKF_3DR is
43 maintained more than in the EnKF. Furthermore, it is found that the EnKF_3DR outperforms the
44 EnKF.

48 **1. Introduction**

49 Ensemble generation methods seek to create a set of initial perturbations representative of
50 analysis errors in a numerical weather prediction system with the goal to improve its probabilistic
51 forecast performance. The analysis errors can be decomposed into nongrowing and growing
52 (Toth and Kalnay, 1997). The nongrowing errors have large dimensional subspace, which cannot
53 be sampled with a limited number of ensemble members and these errors will typically lose their
54 amplitude rapidly. The growing errors amplify fast and dominate the short-range forecast error
55 growth. Therefore, the success of an ensemble generation method lies on how well its
56 perturbations sample the growing errors in the analysis.

57 The Breeding Vector (BV) method (Toth and Kalnay, 1993, 1997) creates perturbations that
58 grow fast by inserting (“breeding”) rescaled errors from previous cycles. After several cycles, the
59 growing component amplifies, and the nongrowing component is eliminated. However, the BV
60 method alone is insufficient to systematically capture all initial uncertainties (Annan, 2004;
61 Buizza et al., 2005). Therefore, an improved version of the BV, the Ensemble Transform (ET)
62 method is introduced to generate initial perturbations which are globally transformed from the
63 forecast perturbations (Bishop and Toth, 1999; Wei et al., 2008). As the BV method, the ET also
64 generates a flow-dependent spatial structure and is able to represent fast growing component of
65 analysis errors with minimal computer expense. The advantages of the ET are that the
66 perturbations have the maximum number of effective degrees of freedom and are more
67 consistent with the data assimilation system due to their orthogonalization in the inverse analysis
68 error variance norm (Wei et al., 2008); more importantly ET outperforms BV in standard
69 probabilistic skill scores.

70 The analysis error variance decides the analysis perturbation globally during the
71 transformation, but the initial spread distribution can be regionally inconsistent with the analysis
72 error variance due to the limited ensemble size compared with the state dimension (McLay et al.,
73 2008; Wei et al., 2008). McLay et al. (2010) performed the local ET with partitioning the global
74 domain into latitude bands or latitude-longitude blocks, resulting in a better agreement with
75 analysis error variance and improved ensemble performance. At NCEP, a simple remedy that
76 regional rescaling process is imposed into the ET initialization periodically to make the
77 amplitude of initial perturbations vary in accordance with regional changes of analysis
78 uncertainties. This regional rescaling improved both the spread distribution of initial
79 perturbations and most probabilistic scores with respect to the ET without rescaling (Wei et al.
80 2008). The regional rescaling factor is designed as the ratio of the mask and the square root of a
81 special norm of analysis perturbations at each grid point. The choice of mask is the key to the
82 regional rescaling. In the NCEP operation, the mask is calculated using a long term averaged
83 root-mean-square of analysis error variance in the kinetic energy norm at the 500hPa level
84 obtained from variational data assimilation system (Szunyogh and Toth, 2002; Wei et al., 2008).
85 But the current mask is not adequate enough for using in the context of ensemble forecast system
86 to represent the analysis uncertainties. First and foremost, the two-dimensional (2D) mask cannot
87 represent the vertical structure of analysis uncertainties. To compensate for the underestimate of
88 analysis error, extra inflations with empirical factor have to be applied to the mask for levels
89 below 500hPa at NCEP. But it is obviously not optimal for regional rescaling. Second, the mask
90 was computed from a past decade climatological data, during which the density and accuracy of
91 observation, as well as the data assimilation technology all have greatly changed. Thus, there is a
92 need to update the mask with the estimation of analysis errors from the current real time data

93 assimilation system to make the initial perturbations more consistent with the observations and
94 data assimilation system. As found in the study of Wei et al. (2008), compared with the ET
95 method, the ETR failed to show high spread in the southern-ocean storm track area due to the
96 mask, which indicated a more accurate time-dependent mask was necessary. Third, the total
97 energy norm may be more reasonable to measure the magnitude of initial perturbations than the
98 kinetic energy norm. Palmer et al. (1998) found that the total energy is more consistent with
99 analysis error statistics than the streamfunction, enstrophy or kinetic energy metric. Some
100 previous studies also considered parts of these problems above and designed different masks.
101 Wang and Bishop (2003) chose the square root of the seasonally and vertically averaged initial
102 ensemble wind variance from the Ensemble Transform Kalman Filter (ETKF) ensemble as the
103 mask applied in the BV method. Magnusson et al. (2008) designed vertically integrated
104 estimation of analysis errors using total energy norm from four-dimensional variational (4D-Var)
105 assimilation system as the mask. In this paper, a new mask will be defined by 3D analysis
106 uncertainty measured in total energy norm obtained from the 80-member ensemble analysis
107 generated by the NCEP's hybrid 3D-Var/EnKF system (Wang et al., 2013). The sensitivity of
108 ETR perturbations and forecast skill to the mask in the NCEP GEFS will be explored.

109 Relative to the variational data assimilation method with static background error, the
110 ensemble-based data assimilation has the ability to provide flow-dependent estimates of the
111 background error. Moreover, the ensemble-based data assimilation generates a set of initial
112 analysis to initialize the ensemble of predictions in the next cycle and also to provide an estimate
113 of the analysis error, which unifies the ensemble forecast and data assimilation steps.
114 Consequently, many Numerical Weather Prediction (NWP) centers are adopting the use of
115 ensemble technology to produce analysis in the data assimilation system and initial conditions in

116 the ensemble prediction system simultaneously, such as the Meteorological Service of Canada
117 (MSC) and the European Centre for Medium-Range Weather Forecasts (ECMWF). Researches
118 demonstrate that the ensemble-based data assimilation is beneficial for both systems by
119 producing flow-dependent estimates of analysis uncertainty and background error uncertainty
120 (Buehner et al., 2010a, b; Buizza et al., 2008, 2010). A hybrid 3D-Var/EnKF data assimilation
121 system became operational on 22 May 2012 at the NCEP (Wang et al., 2013). In this system, the
122 background error is created by a combination of static background error from the 3D-Var and
123 flow-dependent background error produced from the EnKF, and the EnKF perturbations are
124 recentered on the hybrid analysis. The hybrid 3D-Var/EnKF provided better analyses and
125 subsequent forecasts than the previous operational 3D-Var (at
126 <http://www.emc.ncep.noaa.gov/GFS/impl.php>).

127 Since the ETR method is able to maximize the effective degrees of perturbation freedom
128 without extra cost of computer resources, applying the ETR on other ensemble analysis (e.g.
129 multi-center analysis, or analysis from ensemble-based data assimilation) may have a positive
130 impact on the quality of initial conditions. The availability of the EnKF in the NCEP Global Data
131 Assimilation System (GDAS) provides alternative ensemble initial conditions for the operational
132 GEFS. The performance of the EnKF and ETR perturbations in the NCEP operational
133 environment is compared and will be presented in another paper. In this study, the EnKF
134 ensemble analysis will be transformed and rescaled by ET_3DR, and the impact will be explored.

135 In the next section, the methodology of ET_3DR and EnKF_3DR are described. Section 3
136 investigates the horizontal and vertical distributions of perturbations generated by the ETR and
137 ET_3DR, and compares their forecast performances. In section 4, the characteristic of initial

138 perturbations generated by the EnKF and EnKF_3DR is analyzed, and their forecast skills are
 139 compared. The conclusions are summarized in section 5.

140 **2. Initialization methodologies and experimental design**

141 *2.1 Initialization methodologies*

142 *a. The ET_3DR method*

143 In the ETR scheme (Wei et al., 2006; Wei et al., 2008), the analysis perturbations matrix \mathbf{X}^a
 144 are generated from the forecast perturbations matrix \mathbf{X}^f through an ensemble transformation
 145 matrix \mathbf{T} as follows

146
$$\mathbf{X}^a = \mathbf{X}^f \mathbf{T}, \tag{1}$$

147 where n analysis perturbations \mathbf{x}_i^a ($i=1, 2, \dots, n$) are listed as columns in the matrix \mathbf{X}^a , and n
 148 forecast perturbations \mathbf{x}_i^f ($i=1, 2, \dots, n$) are listed as columns in the matrix \mathbf{X}^f . After the
 149 transformation, all perturbations are orthogonal. As shown in Wei et al. (2008) and Ma et al.
 150 (2012), the transformation matrix \mathbf{T} is given by

151
$$\mathbf{T} = \mathbf{C} \mathbf{\Gamma}^{-1/2}, \tag{2}$$

152 where columns of the matrix \mathbf{C} contain the orthonormal eigenvectors ($c_i, i=1, 2, \dots, n$) of the
 153 matrix $\frac{1}{n-1} (\mathbf{X}^f)^T (\mathbf{P}^a)^{-1} \mathbf{X}^f$, and the diagonal matrix $\mathbf{\Gamma}$ contains the corresponding eigenvalues

154 ($\lambda_i, i=1, 2, \dots, n$), in which the first $n-1$ eigenvalues are non-zero and the last eigenvalue is zero.

155 A diagonal matrix \mathbf{F} is defined by setting the zero eigenvalue in $\mathbf{\Gamma}$ to a non-zero constant α .

156 The diagonal matrix \mathbf{P}^a contains the analysis error variances. Then, a simplex transformation is
 157 performed to mask analysis perturbations be centered on the analysis, but perturbations become
 158 quasi-orthogonal at this step.

159 To make the amplitude of initial perturbations vary in accordance with regional changes of
 160 analysis uncertainties, \mathbf{X}^a is rescaled using a rescaling factor γ which is designed as

$$161 \quad \gamma = \begin{cases} \frac{mask}{pertb}, & \text{if } mask < pertb \\ 1, & \text{if } mask \geq pertb \end{cases}. \quad (3)$$

162 Here, *mask* denotes a long term averaged root-mean-square of analysis error variance; *pertb* is
 163 the square root of a special norm from \mathbf{X}^a at each grid point. If the ratio is larger than 1.0, the
 164 rescaling factor will be set to 1.0, which means the perturbations can grow freely; otherwise the
 165 amplitude will be rescaled to the size of the mask.

166 The mask used in the current NCEP operational GEFS is a 2D mask, which is computed from
 167 a long term averaged root-mean-square of analysis error variance in the kinetic energy norm at
 168 the 500hPa level obtained from variational data assimilation system (Szunyogh and Toth, 2002;
 169 Wei et al., 2008).

170 As discussed in section 1, an ensemble of analyses can be obtained from the NCEP operational
 171 hybrid 3D-Var/EnKF data assimilation system directly, which can provide a flow-dependent
 172 estimate of analysis error. In this study, the 3D mask is defined by the root-mean-square of the
 173 deviation total energy norm TE computed from 80-member EnKF analysis

$$174 \quad TE = \frac{1}{80} \sum_{i=1}^{80} \sqrt{\frac{1}{2}(u_i'^2 + v_i'^2 + \kappa T_i'^2)}, \quad (4)$$

175 where u_i' , v_i' and T_i' ($i=1, 2, \dots, 80$) are the deviation of the i th EnKF member from EnKF mean

176 analysis for the wind components and temperature. $\kappa = \frac{c_p}{T_r}$ equals approximately $4.0JKg^{-1}K^{-2}$ in

177 which c_p is the special heat at constant pressure and T_r is the reference temperature. For the

178 purpose of representative of the typical large-scale components and considering the most recent

179 behavior of analysis uncertainties meanwhile, the decaying average method (Cui et al., 2012) is
180 employed to accumulate the mask, given by

$$181 \quad TE_{ave}(t) = (1-w)TE_{ave}(t-1) + wTE(t). \quad (5)$$

182 Here, the averaged mask $TE_{ave}(t)$ is updated by the prior period averaged mask $TE_{ave}(t-1)$ and
183 the most recent $TE(t)$ with the weight coefficient w ($w=2\%$ in this study). To preserve most of
184 the dynamical balance in the perturbations, the mask is smoothed horizontally with a spectral
185 filter.

186 Figure 1a, b show the horizontal distribution of the 2D and 3D mask at 500hPa over the period
187 1 September - 30 November 2012. The 2D mask obtained from static analysis error estimation,
188 has large amplitude over the poorly observed oceans and small amplitude over the data-rich
189 continents. Over the mid-latitudes, the 3D mask estimated from the EnKF data assimilation
190 system has relatively small amplitude over the continents compared to the oceans, but this
191 property is not quite obvious like the 2D mask. That is because the EnKF ensemble analysis used
192 to produce the 3D mask is processed with a multiplicative inflation algorithm to account for
193 unrepresented error sources during the generation (Whitaker and Hamill, 2012). The values of
194 inflation are proportional to the amount observations reduce the ensemble spread, which are
195 large in regions of dense observations. The analysis error should be not only associated with the
196 observation network but also the distribution of the atmospheric instability (Hamill et al., 2003).
197 The 3D mask is more flow-dependent relative to the 2D mask. For example, over the northern
198 and southern extra-tropics, for the 3D mask we see that areas over the maximum amplitude are
199 respective around 60°N and 60°S corresponding to main regions of baroclinic energy conversions,
200 but the maximum areas are over the poles in the 2D mask. That may solve the problem that the
201 old rescaling factor cannot reduce the amplitudes enough at the higher latitudes (Toth and

202 Kalnay, 1997). Another striking difference is located in the tropics, which will be discussed in
203 the next section. Figure 2 shows the vertical profile of the 2D and 3D mask over the same period
204 as Fig. 1. In the 3D mask, the amplitude increases with the altitude and decreases after it reaches
205 the maxima between the 300hPa and 100hPa. The vertical structure cannot be represented with
206 the 2D mask.

207 *b. The EnKF_3DR method*

208 As illustrated in the Fig. 3, the following steps are performed to initialize the ensemble with
209 the EnKF_3DR method. Firstly, use the EnKF method (Whitaker and Hamill, 2002) to generate
210 ensemble analysis. In this study, the 80 EnKF analyses are directly obtained from the NCEP
211 hybrid 3D-Var/EnKF data assimilation system. Secondly, compute the ensemble mean analysis
212 and the 80 ensemble deviations from the ensemble mean analysis. The root-mean-square of the
213 deviation total energy norm is calculated by Eq. (4). Finally, apply ET_3DR method onto the 80
214 EnKF perturbations to generate 80 EnKF_3DR perturbations. The rescaling mask used in the
215 ET_3DR method is the estimate of analysis error computed at the second step.

216 *2.2 Experimental design*

217 Four sets of ensemble generation experiments (ETR, ET_3DR, EnKF and EnKF_3DR) are
218 performed using the NCEP Global Forecast System (GFS) model with a T254 horizontal
219 resolution, 42 sigma-p hybrid vertical levels. The analysis is truncated from the T574L64
220 analysis provided by the NCEP GDAS. The initial perturbations for the ETR and EnKF
221 experiments are obtained from the operational GEFS and hybrid 3D-Var/EnKF data assimilation
222 system respectively. The methods used to generate the ET_3DR and EnKF_3DR ensemble initial
223 perturbations are described in subsection 2.1. The perturbations are updated every 6-hour for 80-
224 member and only 20-member is chosen for long forecasts due to limited computational resources.

225 The ET_3DR initial perturbation cycles are performed from 1 September to 30 November 2012
226 and the first 10 days are used for the system to spin-up. The 8-day long forecasts of the four sets
227 experiments are produced once per day (00 UTC) between 11 September and 30 November 2012
228 (81 cases). To represent model error, all experiments use the Stochastic Total Tendency
229 Perturbation (STTP) (Hou et al., 2006, 2008) as in the NCEP operational GEFS. Verification
230 results are presented for 500hPa geopotential height (Z500); 850hPa temperature (T850); and
231 250hPa, 850hPa, 10m u -components of wind (U250, U850, and U10m) over the extra-tropics of
232 the Northern Hemisphere (NH, 20°-80°N), the extra-tropics of the Southern Hemisphere (SH,
233 20°-80°S) and the Tropics (TR, 20°S-20°N).

234 **3. ETR versus ET_3DR**

235 *3.1 Initial perturbation distribution*

236 Figure 4 shows the vertical profile of the square root of total energy of perturbations at
237 different lead times for the ETR and ET_3DR experiments. Over the NH, the ETR has larger
238 initial amplitude compared to the ET_3DR at the lower levels and the other maximum of initial
239 perturbations is at 250hPa, which is slightly smaller than the ET_3DR (the left panel of Fig. 4a).
240 The left panel of Figs. 4b, c, d shows that the ET_3DR grows faster than the ETR. After 12-h,
241 the amplitude of ET_3DR perturbations gets closer to the ETR below 700hPa, and the difference
242 becomes larger than that at the initial time above 700hPa. After 48-h, the perturbations of the
243 ET_3DR are larger than the ETR for all levels. Over the SH (the middle panel of Figs. 4a, b, c,
244 d), the situation exhibits similarity with that over the NH. The growth rate over the TR (the right
245 panel of Figs. 4a, b, c, d) is lower than that over the NH and SH. At the upper levels, the
246 maximum of the ETR (ET_3DR) initial perturbations is at 200hPa (100hPa) and their growth
247 rates are comparable. At the lower levels, to compensate for the slow growth of the ETR

248 perturbations, the amplitude of its initial perturbations is much larger than that for the ET_3DR,
249 but the amplitude of perturbations for the ETR is still caught up by the ET_3DR after 96-h. The
250 fast growth of the ET_3DR perturbations compared to the ETR shown in Fig. 4 may be due to its
251 structure of initial perturbations which could better sample the analysis error with applying the
252 3D regional rescaling mask. To further illustrate the details of the initial perturbations, the
253 horizontal and vertical distributions will be analyzed below.

254 Figure 5 shows the horizontal distribution of the square root of total energy of initial
255 perturbations on the 500hPa level for the two experiments. It is found that the regional rescaling
256 masks applied for the ETR and ET_3DR experiments have great impact on their initial
257 perturbations. For the ETR (Fig. 5a), the maxima are around the poles. In addition, the dominant
258 feature of initial perturbations is the locations of the maxima and minima coinciding with the
259 distributions of oceans and continents. The maxima are over the North Pacific, Atlantic and
260 Indian Oceans, and the minima are located in the North America, Eurasia and Australia. For the
261 ET_3DR (Fig. 5b), the large initial perturbations are over the meridional bands around 60°N,
262 60°S and the equator, which seem to be related to the baroclinic zones and tropical convection
263 zones respectively. The flow-dependent ET_3DR initial perturbations will be beneficial for
264 obtaining a sufficient dispersed ensemble in the medium range.

265 To investigate the connection between the initial perturbations and baroclinic instability, the
266 correlation coefficients between the Eady index and the square root of total energy of initial
267 perturbations over the NH and SH for both experiments are shown in Fig. 6 in which the shades
268 indicate that the correlation is statistically significant at the 95% confidence interval. The Eady
269 index which is a simple measure of the most unstable Eady mode, is defined as (Hoskins and
270 Valds, 1990)

271
$$\sigma_E = 0.31 \frac{f}{N} \frac{du}{dz}, \quad (6)$$

272 where f is the Coriolis parameter, N is the static stability, u is the magnitude of the vector wind.

273 Here, $N = (g \frac{d \ln \theta}{dz})^{\frac{1}{2}}$ and $\frac{du}{dz}$ are computed using the 300hPa and 1000hPa potential temperature

274 and wind from the NCEP Final Analyses. Over the NH (Figs. 6a, b), the areas of initial

275 perturbations for the ET_3DR which are statistically significantly correlated with the Eady index

276 are larger than the ones for the ETR. Especially, the correlation coefficients are higher than 0.6

277 even up to 0.8 over the western part of the Pacific Ocean and the Atlantic Ocean in the ET_3DR

278 experiment. Over the SH (Figs. 6c, d), the correlations in the two experiments are low compared

279 to those over the NH.

280 Over the TR, the deep convection has an important role on the development of perturbations

281 through the release of latent heating. To illustrate the relationship between the initial

282 perturbations and deep convection, the outgoing longwave radiation (OLR) which is a common

283 measure of the intensity of the tropical convection is plotted in Fig. 7. The low value of OLR

284 represents intense tropical convection. For the ET_3DR (Figs. 5b, 7), the locations of the

285 maxima of initial perturbations accurately coincide with the intense deep convection zones (low

286 OLR) except the maximum over the eastern Pacific Ocean. This connection cannot be detected at

287 all in the ETR experiment (Figs. 5a, 7).

288 Figure 8 shows the zonal average of the square root of total energy of initial perturbations. For

289 the ETR (Fig. 8a), below the 200hPa level, the minima of initial perturbations on both

290 hemispheres are around 60°N and 40°S, respectively. Over the TR, there are two maxima at

291 10°N, 300hPa and 950hPa. Above the 200hPa level, the perturbations decrease with the height

292 over the globe. For the ET_3DR (Fig. 8b), the amplitude is slightly larger over the SH than over

293 the NH. The maxima are around 55°N and 55°S at 300hPa, which correspond with the
294 subtropical jet regions. Over the TR, the maximum is around 10°N, 100hPa, near the tropical
295 easterlies jet region.

296 *3.2 Ensemble forecast skill*

297 The verification methods used to evaluate the ensemble forecast skills with ETR and ET_3DR
298 initial perturbations include Root Mean Square Error (RMSE; Toth et al., 2003) of ensemble
299 mean and Continuous Ranked Probability Score (CRPS; Toth et al., 2003; Wilks, 2006). The
300 paired block bootstrap algorithm (Hamill, 1999) is used to estimate the statistical significance of
301 differences in scores. In this study, 95% confidence interval is computed from a bootstrap
302 resampling using 1000 random samples of the 81 cases.

303 *a. RMSE and ensemble spread*

304 Figures 9a, b, c, d show the ensemble mean RMSE and ensemble spread for U250, Z500,
305 U850, and T850 over the NH. Comparing the RMSE of the ETR and ET_3DR experiments, the
306 results have no significant differences for all lead times. Regarding the ensemble spread, there
307 are substantial differences between the two experiments. For U250, the ET_3DR and ETR have
308 the same size of initial perturbations, but the ET_3DR grows faster than the ETR and keeps
309 being consistent with the RMSE for all lead times as a perfect ensemble forecast system should
310 do (Fig. 9a). For the indirect model variable Z500, figure 9b shows that the ET_3DR starts from
311 a larger spread and overestimates the ensemble mean errors, but the amplitude of initial
312 perturbations could be tuned further to give a similar spread to the errors at the initial time. For
313 U850 and T850, the ET_3DR initial perturbations are much smaller than the ETR, but the spread
314 catches up after 24-h and gets close to the RMSE gradually (Figs. 9c, d). Over the SH (Figs. 10a,
315 b, c, d), the results are similar to the ones over the NH.

316 Figures 11a, b show the RMSE and spread for U850 and U10m since the wind field is of more
317 interest than the mass field over the TR. The ETR starts from a much larger spread than the
318 ET_3DR and decays during the first 2-d. The spread of ET_3DR has higher growth rate than the
319 ETR especially for the first 2-d. As found in section 3.1, that is attributed to the close connection
320 between the initial perturbations of the ET_3DR over the TR and the tropical deep convection
321 which is an important factor for the development of perturbations. The ETR has significantly
322 higher RMSE than the ET_3DR at 12-h lead time. Both experiments produce smaller spread than
323 the RMSE. Because the growth in the ensemble spread over the TR is mostly determined by the
324 physical processes, while that over the NH and SH is mainly influenced by the dynamic
325 instability, sampling the model related errors plays a more important role on the ensemble spread
326 over the TR.

327 Overall, the main advantage of the ET_3DR as the figures shown is the higher growth rate
328 unlike the ETR method which artificially increases the amplitude of the initial perturbations to
329 compensate for the low growth rate of spread with the cost of negative effects on the
330 performance at short lead times. This advantage is especially obvious at lower levels and over
331 the TR.

332 *b. Continuous Ranked Probability Score (CRPS)*

333 The CRPS is used to measure the reliability and resolution of ensemble-based probabilistic
334 forecasts by calculating the distance between the predicted and the observed cumulative
335 distribution functions of scalar variables. The smaller the score is, the better the quality of the
336 probabilistic forecast is. Over the NH, the CRPS for U250 is similar for the two experiments (Fig.
337 12a). The ETR has significantly smaller score than the ET_3DR for the first 12-h for Z500 (Fig.
338 12b). There are more improvements on the probabilistic forecast score for lower levels compared

339 to upper levels using the ET_3DR initial perturbations. For U850, the ET_3DR produces
340 statistically significantly better probabilistic forecast for the first 4-d than the ETR (Fig. 12c). For
341 T850, the ET_3DR has slightly but statistically significantly better performance for the first 2-d
342 than the ETR (Fig. 12d). Over the SH (Figs. 13a, b, c, d), the results are generally similar to that
343 over the NH, except that the ET_3DR presents statistically significantly smaller value than the
344 ETR only for lead times up to 12-h for U850 and T850. Over the TR (Fig. 14), the ET_3DR has
345 statistically significantly better performance than the ETR for almost all 8-d lead times except
346 day 1.5-2.5 for U850. For U10m, the ET_3DR has statistically significant advantage over the
347 ETR for the first 1-d.

348 **4 EnKF versus EnKF_3DR**

349 *4.1 Initial perturbation*

350 Initial perturbations should span as many unstable directions of the atmosphere as possible
351 with limited ensemble members. The eigenvalue spectra of the covariance matrix of the initial
352 perturbations can be used to evaluate the distribution of the amounts in independent directions.
353 Figure 15 shows the mean eigenvalue spectra for Z500 during the period 11 September - 30
354 November 2012 over the globe. It is found that the initial perturbations of the EnKF are too
355 much contained in the direction of the first mode. The EnKF_3DR has flatter spectra than the
356 EnKF, implying that the ensemble members are more independent than the EnKF, which may
357 have potentially positive impact on the ensemble performance.

358 *4.2 Ensemble forecast skill*

359 The results of the EnKF and EnKF_3DR experiments will be compared in this section using
360 the same verification methods as in section 3.2.

361 *a. RMSE and ensemble spread*

362 In Fig. 16, the RMSE and ensemble spread for U250, Z500, U850 and T850 over the NH are
363 shown. Comparing the RMSE, the EnKF_3DR is slightly better than the EnKF for U250 and
364 Z500 (Figs. 16a, b), but the difference is not statistically significant for U250 and only
365 significant for the first 1-d for Z500. Results for U850 and T850 (Figs. 16c, d) show that the
366 EnKF_3DR has significantly smaller RMSE than the EnKF for the first 3.5-d. Regarding the
367 ensemble spread, the growth rates are basically similar for the two experiments. For U250 and
368 Z500, the initial spread for the EnKF_3DR is slightly smaller than the EnKF, and the spread for
369 the EnKF_3DR is more consistent with the RMSE compared to the EnKF until 4-d for U250 and
370 all lead times for Z500 (Figs. 16a, b). For U850 and T850, the spread grows somewhat slower
371 than the RMSE for short lead times, but becomes almost equal to the RMSE with the increasing
372 of forecast length (Figs. 16c, d).

373 Over the SH, the difference of the RMSE between the two experiments for U250 is also not
374 significant for all lead times (Fig. 17a). The forecast length, during which the EnKF_3DR has
375 significantly smaller RMSE than the EnKF, extends to 6.5-d for Z500 (Fig. 17b). For U850 and
376 T850, the EnKF_3DR produces smaller RMSE than the EnKF, and the difference is statistically
377 significant for all lead times (Figs. 17c, d). The spread for U250 and Z500 in the EnKF_3DR
378 experiment is more consistent with the RMSE compared to the one in the EnKF experiment
379 which is much larger than the RMSE (Figs. 17a, b). Similar to the one over the NH, the spread
380 for U850 and T850 grows slower than the RMSE during the first 3 to 4-d, and then becomes
381 almost equal to the RMSE with the increasing of forecast length.

382 Results for the TR (Fig. 18) show that U850 and U10m for both experiments appear to
383 produce much less spread than the RMSE due to the under-sampling of model related errors. The

384 spread for the EnKF_3DR grows slightly slower than the EnKF, but the RMSE is substantially
385 smaller than the EnKF.

386 *b. CRPS*

387 The CRPS for U250 shows similar scores between the two experiments for both hemispheres
388 and only differs significantly for the first 12-h over the SH (Figs. 19a and 20a). For Z500, the
389 EnKF_3DR produces slightly better probabilistic forecast than the EnKF over the NH, and the
390 difference is significant for up to 2-d (Fig. 19b). Over the SH, the improvement becomes more
391 apparent that is significant for 6.5-d (Fig. 20b). For U850 and T850, the EnKF_3DR has
392 substantially better performance than the EnKF for both hemispheres. The difference is
393 statistically significant until 4-d over the NH, and all lead times over the SH (Figs. 19c and 20c).
394 Over the TR, for all lead times, the EnKF_3DR shows significantly better score for both U850
395 and U10m (Fig. 21).

396 **5 Conclusions and discussion**

397 In the ETR method, the rescaling mask plays a critical role to constrain the amplitude of initial
398 perturbations to reflect regional changes of analysis error. While the ETR used in the NCEP
399 GEFS has proved to improve the spread and probabilistic skill of the ensemble forecasts over
400 both BV and ET methods, its mask has several limitations, which in this study we attempt to
401 address. There are three main modifications to the mask. First and foremost, for representing the
402 vertical structure of analysis error, the 3D mask is employed instead of the 2D mask. This is the
403 most advantage of the ET_3DR compared to the ETR. In the ETR method, due to the vertical-
404 constant mask used, extra inflations have to be applied to the mask for levels from model bottom
405 to 500hPa with empirical factors to compensate for the underestimate of analysis errors. Second,
406 with the availability of an ensemble of analyses from the hybrid 3DVar-EnKF data assimilation

407 system, on each data assimilation cycle a flow-dependent error variance is computed with real
408 observations, which is associated with both the dynamics of the day and the observation density
409 distribution. This new analysis error variance replaces the static analysis error variance. Third,
410 the kinetic energy norm is changed into the total energy norm to measure the magnitude of initial
411 perturbations. Results with the ETR and ET_3DR experiments performed from 11 September to
412 30 November 2012 using the NCEP GFS indicate that these updates have direct impact on the
413 perturbations. The horizontal distribution of the ETR initial perturbations at 500hPa coincides
414 with the distribution of oceans and continents, but is not consistent with the flow. Because of the
415 flow-dependent mask applied in the ET_3DR, the large amplitudes of the initial perturbations
416 connect better with the areas of baroclinic instability over the NH and the areas of deep
417 convection over the TR, which is beneficial for obtaining a sufficient dispersed ensemble in the
418 medium range. The difference of vertical distribution for the ETR perturbations is small due to
419 the vertical-constant mask, while the maxima of vertical distribution for the ET_3DR
420 perturbations correspond to the subtropical jet region and tropical easterlies jet region. Since the
421 amplitude of the initial perturbations for the ET_3DR is better consistent with the distribution of
422 the atmospheric instability regions, the spread grows much faster than the ETR, especially at the
423 lower levels and over the TR. Consequently, the choice of mask is important to perturbation
424 growth and ensemble performance for the NCEP GEFS.

425 Since the ETR method is able to maximize the effective degrees of perturbation freedom
426 without extra cost of computer resources, with the availability of the EnKF analyses in the NCEP
427 GDAS, the EnKF_3DR method is designed in this study by applying ET_3DR on EnKF
428 ensemble analysis. The eigenvalue spectra of the covariance matrix of the initial perturbations
429 show that the ensemble members of the EnKF_3DR are more independent than the EnKF. By

430 evaluating the ensemble performance, it is found that the EnKF_3DR is substantially better than
431 the EnKF, especially at the lower levels and over the TR.

432 The EnKF may be considered as the potential candidate of the NCEP operational GEFS initial
433 perturbation method in the next implementation. However, from the results of this study, we can
434 find that applying ET_3DR on EnKF is more beneficial for the improvement of the ensemble
435 forecast performance than the EnKF. Further studies will explore the results using this strategy
436 for other seasons. Furthermore, due to the merit of the ETR method, it may be also considered to
437 apply on other ensemble analysis, such as multi-center analysis.

438 Although the results of this study indicate that the improvement of the mask benefits the
439 ensemble performance, this regional rescaling is only a simplified remedy to the complex
440 problem of making initial spread distribution agree with the analysis error variance regionally.
441 Therefore, future research effort should be on practical accounting for all sources of analysis
442 uncertainties.

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447 **References**

448 Annan, J. D., 2004: On the orthogonality of bred vectors. *Mon. Wea. Rev.*, **132**, 843–849.

449 Bishop, C. H., and Z. Toth, 1999: Ensemble transformation and adaptive observations. *J. Atmos.*
450 *Sci.*, **56**, 1748–1765.

451 Buehner, M., P. L. Houtekamer, C. Charette, H. L. Mitchell, and B. He, 2010a: Intercomparison
452 of variational data assimilation and the ensemble Kalman filter for global deterministic NWP.

453 Part I: Description and single-observation experiments. *Mon. Wea. Rev.*, **138**, 1550–1566.

454 —, —, —, —, and —, 2010b: Intercomparison of variational data assimilation and the
455 ensemble Kalman filter for global deterministic NWP. Part II: One-month experiments with
456 real observations. *Mon. Wea. Rev.*, **138**, 1567–158.

457 Buizza, R., P. L. Houtekamer, Z. Toth, G. Pellerin, M. Wei, and Y. Zhu, 2005: A Comparison of
458 the ECMWF, MSC, and NCEP Global Ensemble Prediction Systems. *Mon. Wea. Rev.*, **133**,
459 1076–1097.

460 —, M. Leutbecher, and L. Isaksen, 2008: Potential use of an ensemble of analyses in the
461 ECMWF Ensemble Prediction System. *Quart. J. Roy. Meteor. Soc.*, **134**, 2051–2066.

462 —, —, —, and J. Haseler, 2010: Combined use of EDA- and SV-based perturbations in the
463 EPS. *Newsletter* n. 123, ECMWF, Shinfield Park, Reading RG2-9AX, UK, 22–28.

464 Cui, B., Z. Toth, Y. Zhu, and D. Hou, 2012: Bias Correction for Global Ensemble Forecast. *Wea.*
465 *Forecasting*, **27**, 396–410.

466 Hamill, T. M., 1999: Hypothesis tests for evaluation numerical precipitation forecasts. *Wea.*
467 *Forecasting*, **14**, 155–167.

468 —, C. Snyder, and J. S. Whitaker, 2003: Ensemble forecasts and the properties of flow-
469 dependent analysis-Error covariance singular vectors. *Mon. Wea. Rev.*, **131**, 1741–1758.

470 Hoskins, B. J., and P. J. Valdes, 1990: On the existence of storm-tracks. *J. Atmos. Sci.*, **47**, 1854–
471 1864.

472 Hou, D., Z. Toth, and Y. Zhu, 2006: A stochastic parameterization scheme within NCEP global
473 ensemble forecast system. In Proceedings of the *18th AMS Conference on Probability and*
474 *Statistics*, 29 January –2 February 2006, Atlanta, Georgia.

475 —, —, —, and W. Yang, 2008: Impact of a stochastic perturbation scheme on NCEP global
476 ensemble forecast system. In Proceedings of the *19th AMS Conference on Probability and*
477 *Statistics*, 21–24 January 2008, New Orleans, Louisiana.

478 Houtekamer, P. L., and H. L. Mitchell, 2005: Ensemble Kalman filtering. *Quart. J. Roy. Meteor.*
479 *Soc.*, **131**, 3269–3289.

480 Ma, J., Y. Zhu, R. Wobus, and P. Wang, 2012: An effective configuration of ensemble size and
481 horizontal resolution for NCEP GEFS. *Adv. Atmos. Sci.*, **29**, 782–794.

482 Magnusson L., M. Leutbecher, E. Kallen, 2008: Comparison between singular vectors and
483 breeding vectors as initial perturbations for the ECMWF Ensemble Prediction System. *Mon.*
484 *Wea. Rev.*, **136**, 4092–4104.

485 McLay, J. G., C. H. Bishop, C. A. Reynolds, 2008: Evaluation of the Ensemble Transform
486 Analysis Perturbation Scheme at NRL. *Mon. Wea. Rev.*, **136**, 1093–1108.

487 —, —, and —, 2010: A Local Formulation of the Ensemble Transform (ET) Analysis
488 Perturbation Scheme. *Wea. Forecasting*, **25**, 985–993.

489 Palmer, T. N., R. Gelaro, J. Barkmeijer, and R. Buizza, 1998: Singular vectors, metrics, and
490 adaptive observations. *J. Atmos. Sci.*, **55**, 633–653.

491 Szunyogh, I., and Z. Toth, 2002: The effect of increased horizontal resolution on the NCEP
492 global ensemble mean forecasts. *Mon. Wea. Rev.*, **130**, 1125–1143.

493 Toth, Z., and E. Kalnay, 1993: Ensemble forecasting at NMC: the generation of perturbations.
494 *Bull. Amer. Meteor. Soc.*, **174**, 2317–2330.

495 —, and —, 1997: Ensemble forecasting at NCEP and the breeding method. *Mon. Wea. Rev.*,
496 **125**, 3297–3319.

497 —, O. Talagrand, G. Candille, and Y. Zhu, 2003: Probability and Ensemble Forecasts. *Forecast*
498 *Verification: A practitioner's Guide in Atmospheric Science*, Jolliffe and Stephenson, Wiley,
499 137–163.

500 Wang, X., and C. H. Bishop, 2003: A comparison of breeding and ensemble transform Kalman
501 filter ensemble forecast schemes. *J. Atmos. Sci.*, **60**, 1140–1158.

502 —, D. Parrish, D. Kleist, and J. Whitaker, 2013: GSI 3DVar-based Ensemble-Variational
503 Hybrid Data Assimilation for NCEP Global Forecast System: Single Resolution Experiments.
504 *Mon. Wea. Rev.*, doi:10.1175/MWR-D-12-00141.1, in press.

505 Wei, M., Z. Toth, R. Wobus, Y. Zhu, C. H. Bishop, and X. Wang, 2006: Ensemble Transform
506 Kalman Filter-based ensemble perturbations in an operational global prediction system at
507 NCEP. *Tellus*, **58A**, 28–44.

508 —, —, —, and —, 2008: Initial perturbations based on the ensemble transform (ET)
509 technique in the NCEP global operational forecast system. *Tellus*, **60A**, 62–79.

510 Whitaker, J. S. and T. M. Hamill, 2002: Ensemble data assimilation without perturbed
511 observations. *Mon. Wea. Rev.*, **130**, 1913–1924.

512 —, and —, 2012: Evaluating methods to account for system errors in ensemble data
513 assimilation. *Mon. Wea. Rev.*, **140**, 3078–3089.

514 Wilks, D. S., 2006: *Statistical Methods in the Atmospheric Sciences*. 2nd ed., Academic Press,
515 627pp.

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Figure Caption List

521 Figure 1: The regional rescaling (a) 2D mask and (b) 3D mask at the 500hPa level over the
522 period 1 September - 30 November 2012. The contour interval is 0.2 m s^{-1} .

523 Figure 2: The vertical profile of the 2D mask and 3D mask over the period 1 September - 30
524 November 2012.

525 Figure 3: The flow chart of the EnKF_3DR method.

526 Figure 4: The vertical profile of the square root of total energy (m s^{-1}) of perturbations at (a) the
527 initial time, (b) 12-h, (c) 48-h, and (d) 96-h forecast time for the ETR (dashed) and ET_3DR
528 (solid) experiments at the 500hPa level as a mean for the period 11 September - 30 November
529 2012.

530 Figure 5: The square root of total energy (m s^{-1}) of initial perturbations for the (a) ETR and (b)
531 ET_3DR experiments at the 500hPa level as a mean for the period 11 September - 30
532 November 2012.

533 Figure 6: The correlation coefficients between the Eady index and the square root of total energy
534 of initial perturbations, which are statistically significant at the 95% confidence interval for
535 500hPa level over the period 11 September - 30 November 2012.

536 Figure 7: The average OLR at the 500hPa level over the period 11 September - 30 November
537 2012. The contour interval is 20W/m^2 .

538 Figure 8: The zonal average of the square root of total energy (m s^{-1}) of initial perturbations for
539 the (a) ETR and (b) ET_3DR experiments at the 500hPa level as a mean for the period 11
540 September - 30 November 2012.

541 Figure 9: The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U250, (b)
542 Z500, (c) U850 and (d) T850 over the NH. The vertical bars represent the 95% confidence
543 interval from a paried block bootstrap.

544 Figure 10: Same as in Fig. 9, but over the SH.

545 Figure 11: The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U850 and (b)
546 U10m over the TR for the period 11 September - 30 November 2012. The vertical bars
547 represent the 95% confidence interval from a paried block bootstrap.

548 Figure 12: The CRPS for (a) U250, (b) Z500, (c) U850 and (d) T850 over the NH for the period
549 11 September - 30 November 2012. The vertical bars represent the 95% confidence interval
550 from a paried block bootstrap.

551 Figure 13: Same as in Fig. 12, but over the SH.

552 Figure 14: The CRPS for (a) U850 and (b) U10m over the TR for the period 11 September - 30
553 November 2012. The vertical bars represent the 95% confidence interval from a paried block
554 bootstrap.

555 Figure 15: The mean eigenvalue spectra of the covariance matrix of the initial perturbations for
556 Z500 during the period 11 September - 30 November 2012 over the globe.

557 Figure 16: The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U250, (b)
558 Z500, (c) U850 and (d) T850 over the NH. The vertical bars represent the 95% confidence
559 interval from a paried block bootstrap.

560 Figure 17: Same as in Fig. 16, but over the SH.

561 Figure 18: The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U850 and (b)
562 U10m over the TR for the period 11 September - 30 November 2012. The vertical bars
563 represent the 95% confidence interval from a paried block bootstrap.

564 Figure 19: The CRPS for (a) U250, (b) Z500, (c) U850 and (d) T850 over the NH for the period
565 11 September - 30 November 2012. The vertical bars represent the 95% confidence interval
566 from a paried block bootstrap.

567 Figure 20: Same as in Fig. 19, but over the SH.

568 Figure 21: The CRPS for (a) U850 and (b) U10m over the TR for the period 11 September - 30
569 November 2012. The vertical bars represent the 95% confidence interval from a paried block
570 bootstrap.

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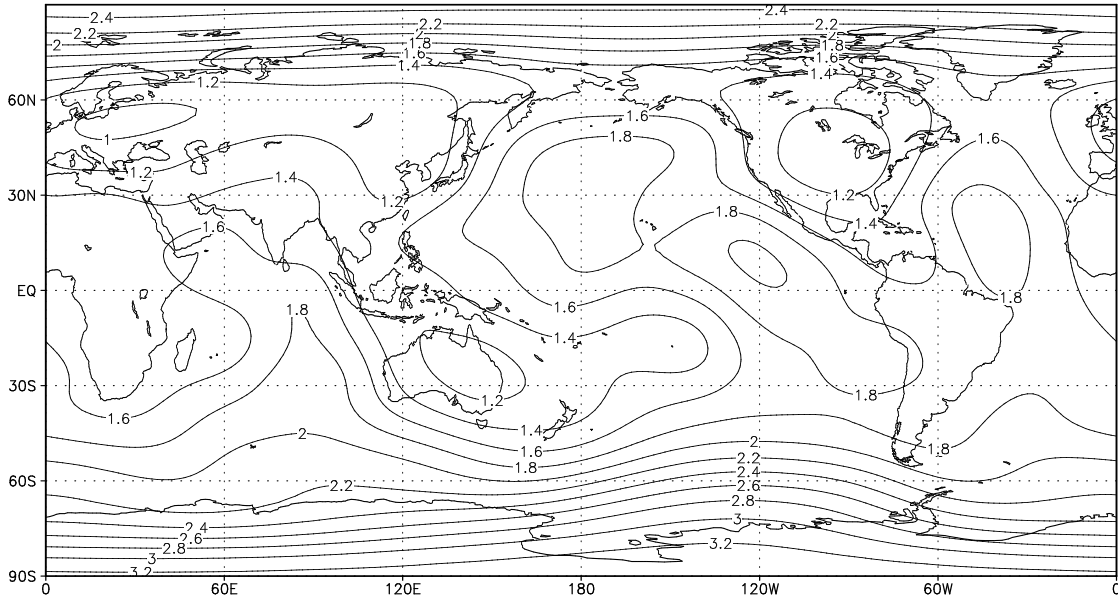
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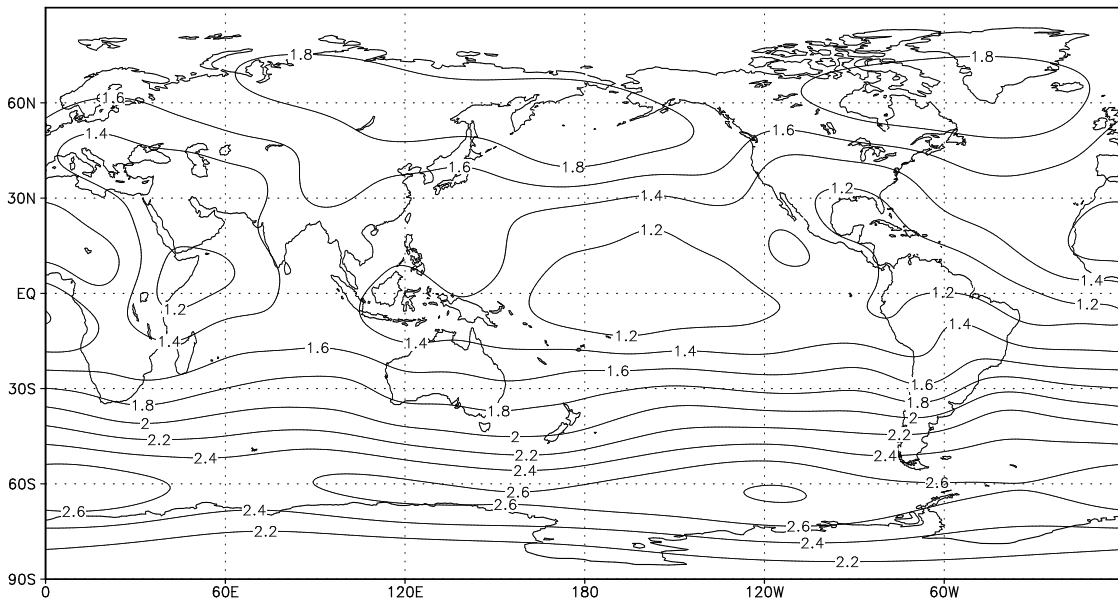
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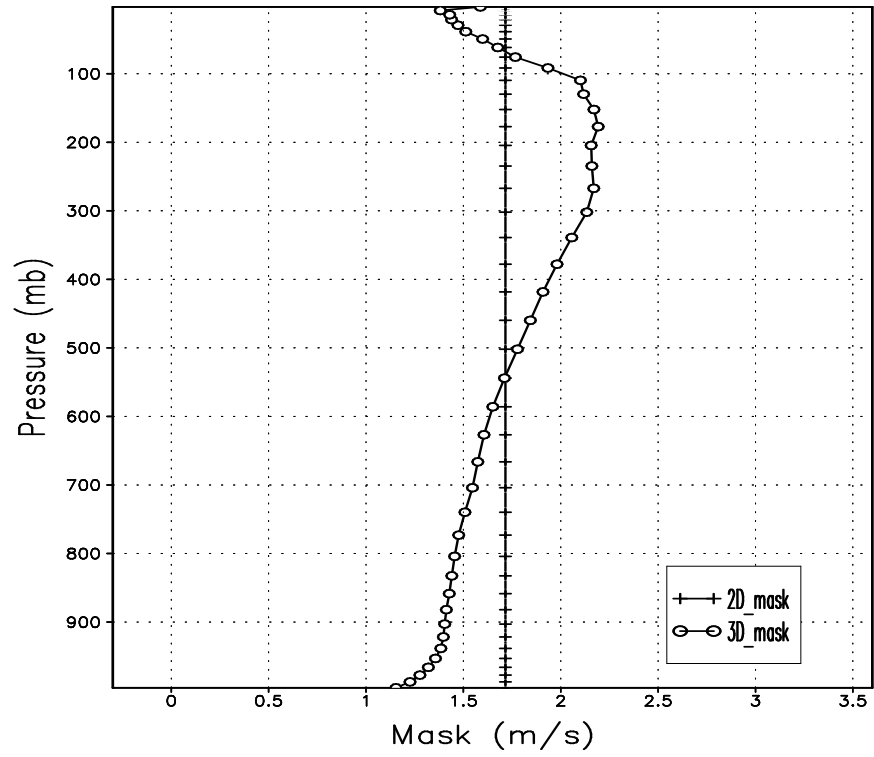
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591 Fig. 1 The regional rescaling (a) 2D mask and (b) 3D mask at the 500hPa level over the period 1

592 September - 30 November 2012. The contour interval is 0.2 m s^{-1} .

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596 Fig. 2 The vertical profile of the 2D mask and 3D mask over the period 1 September - 30

597 November 2012.

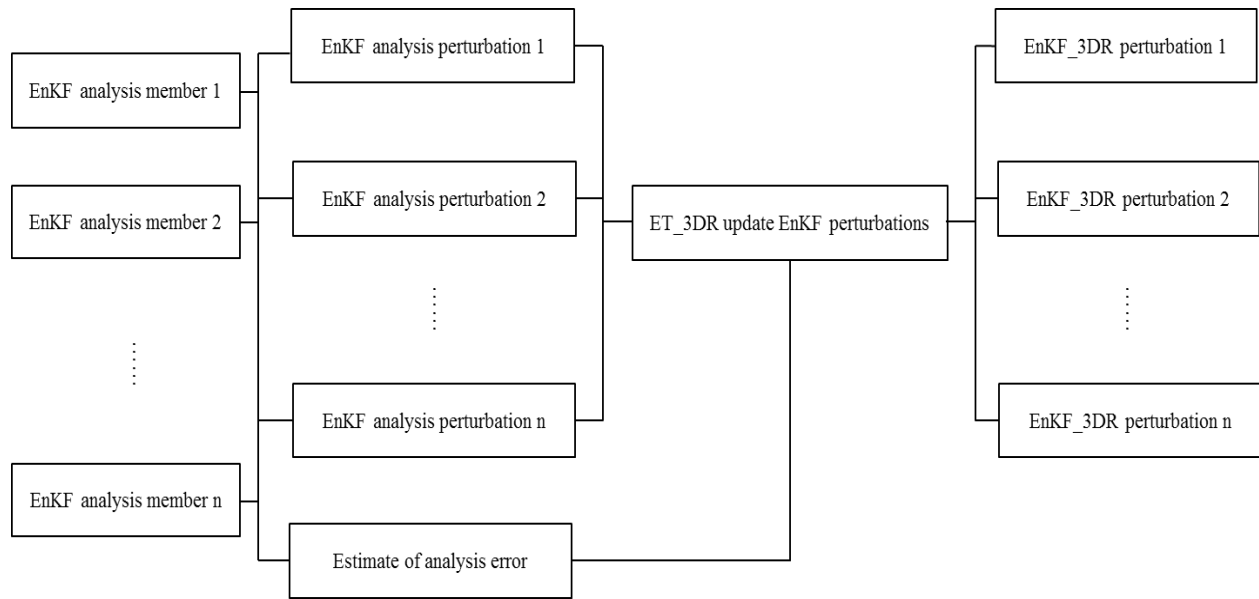
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605 Fig. 3 The flow chart of the EnKF_3DR method.

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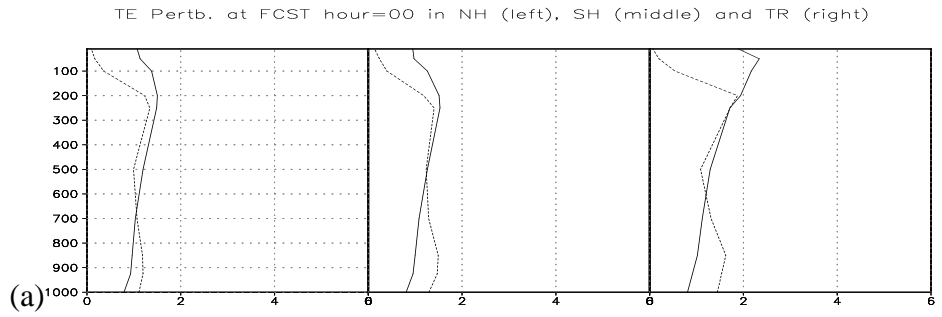
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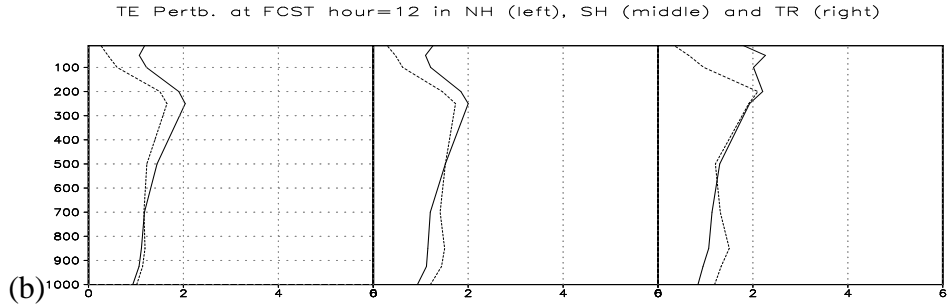
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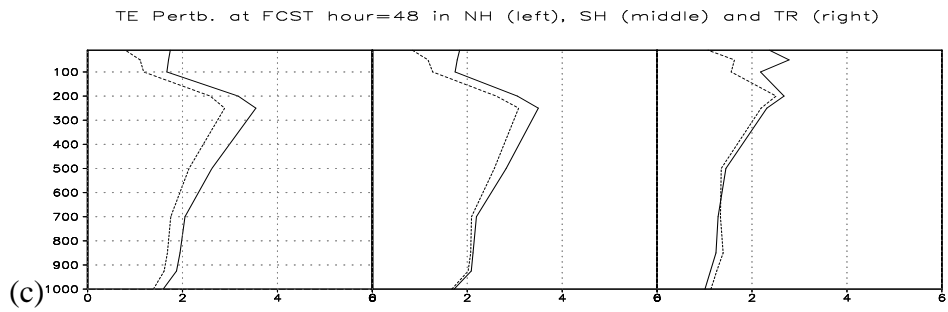
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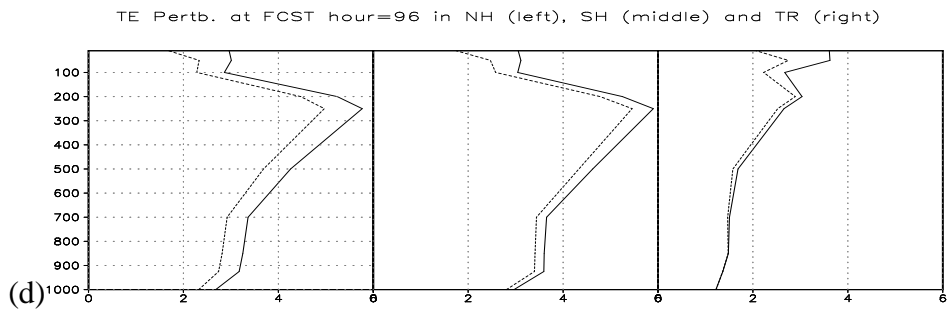
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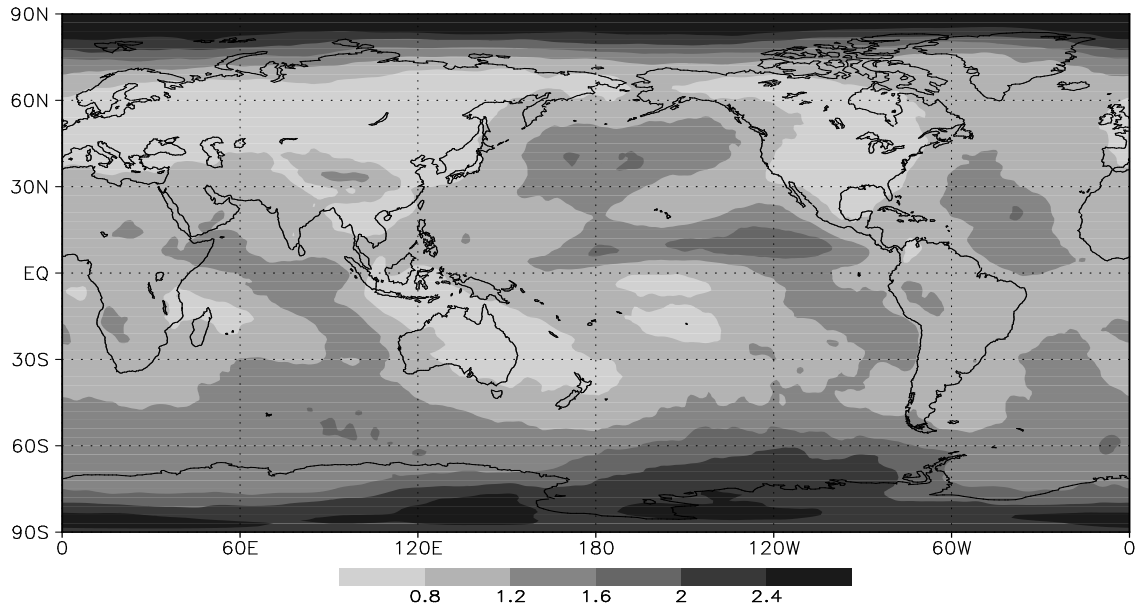
623 Fig. 4 The vertical profile of the square root of total energy (m s^{-1}) of perturbations (a) at the
 624 initial time, after (b) 12-h, (c) 48-h, and (d) 96-h for the ETR (dashed) and ET_3DR (solid)
 625 experiments at the 500hPa level as a mean for the period 11 September - 30 November 2012.

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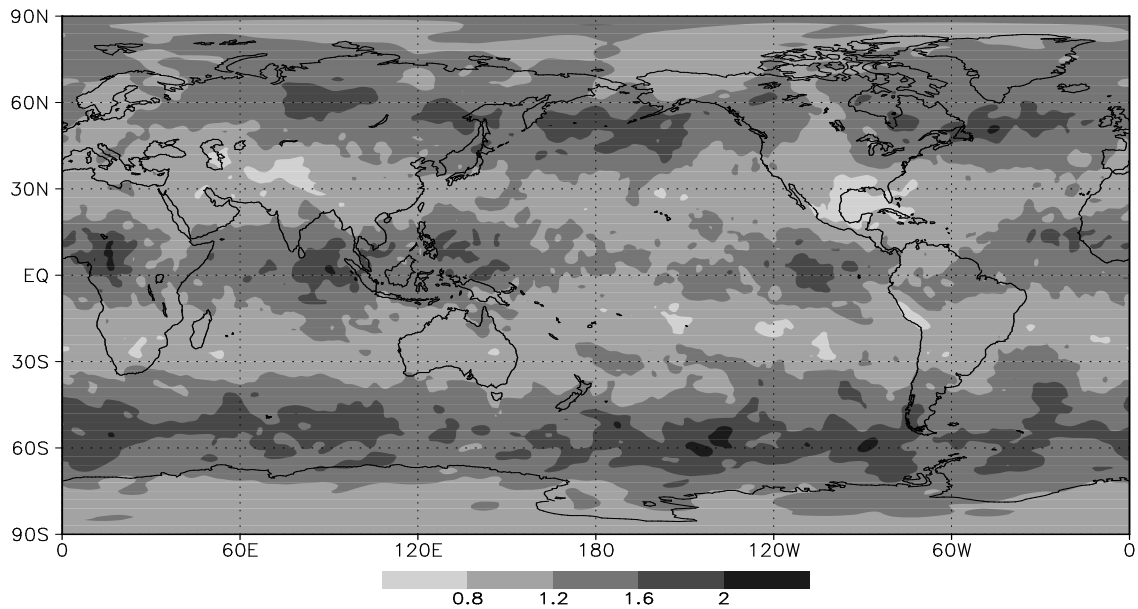
(a) ETR



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(b) ET_3DR



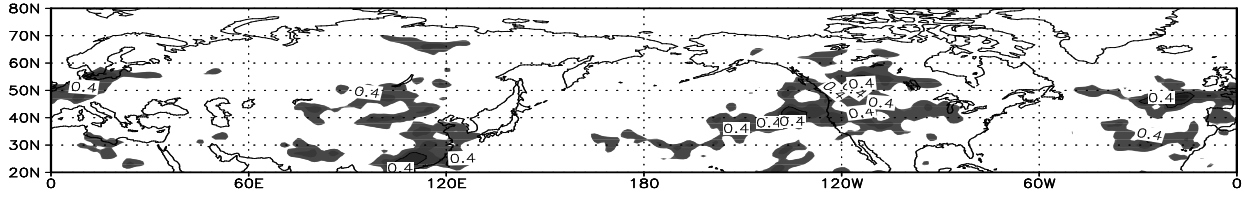
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632 Fig. 5 The square root of total energy (m s^{-1}) of initial perturbations for the (a) ETR and (b)
633 ET_3DR experiments at the 500hPa level as a mean for the period 11 September - 30 November
634 2012.

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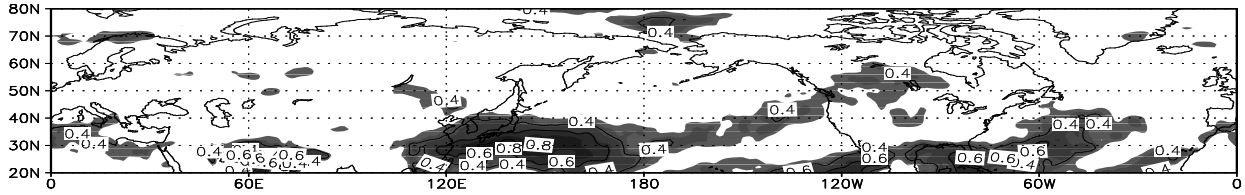
(a) ETR over NH



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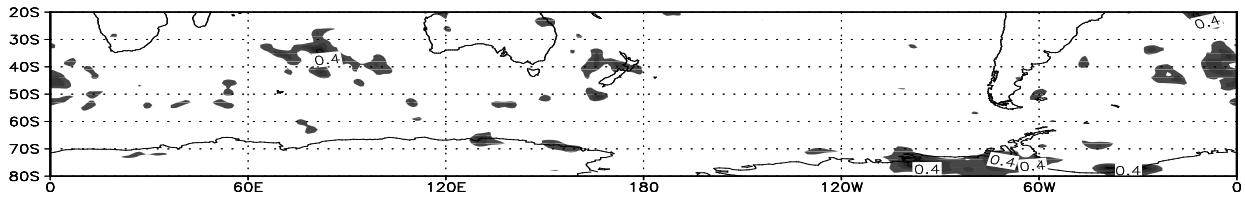
(b) ET_3DR over NH



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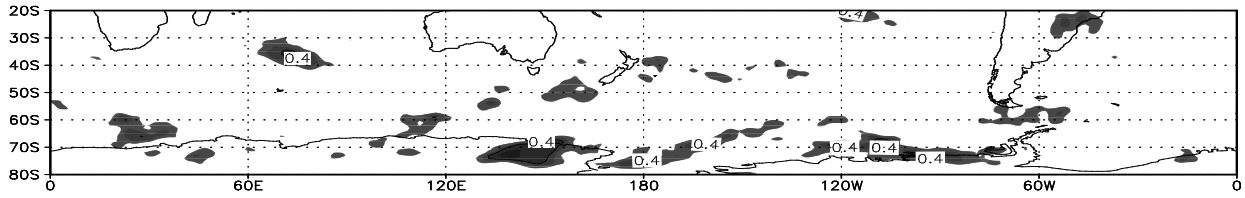
(c) ETR over SH



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(d) ET_3DR over SH



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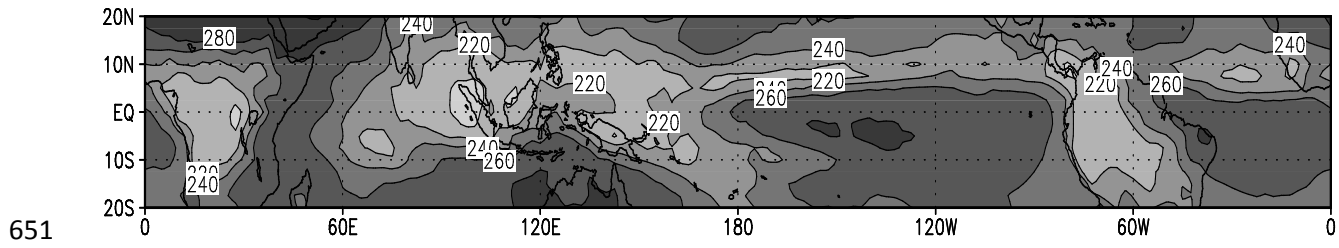
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645 initial perturbations, which are statistically significant at the 95% confidence interval for 500hPa
646 level over the period 11 September - 30 November 2012.

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652 Fig. 7 The average OLR at the 500hPa level over the period 11 September - 30 November 2012.

653 The contour interval is 20W/m².

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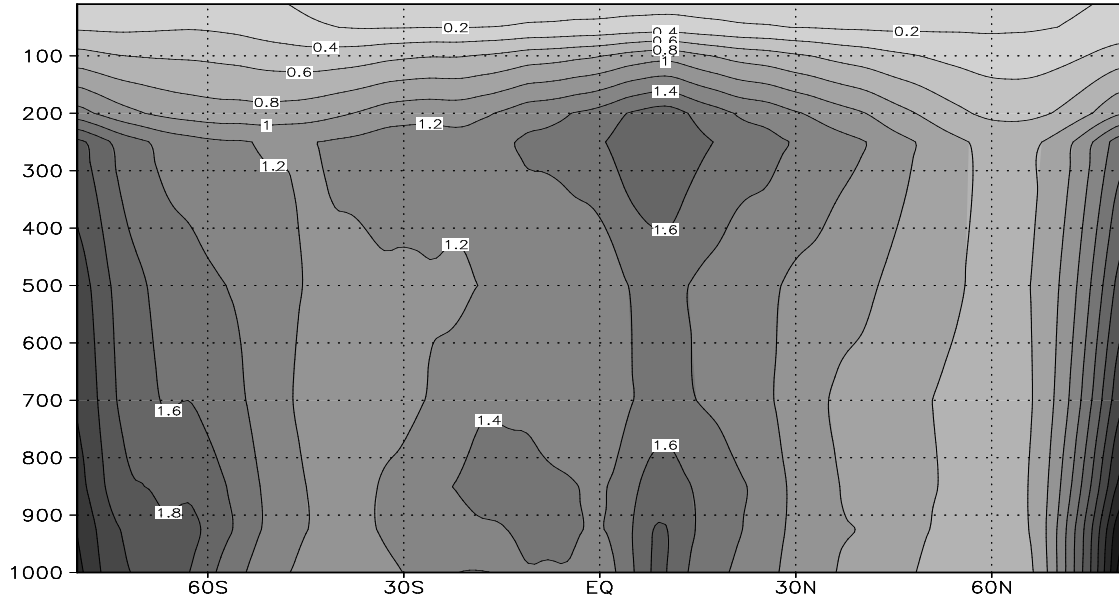
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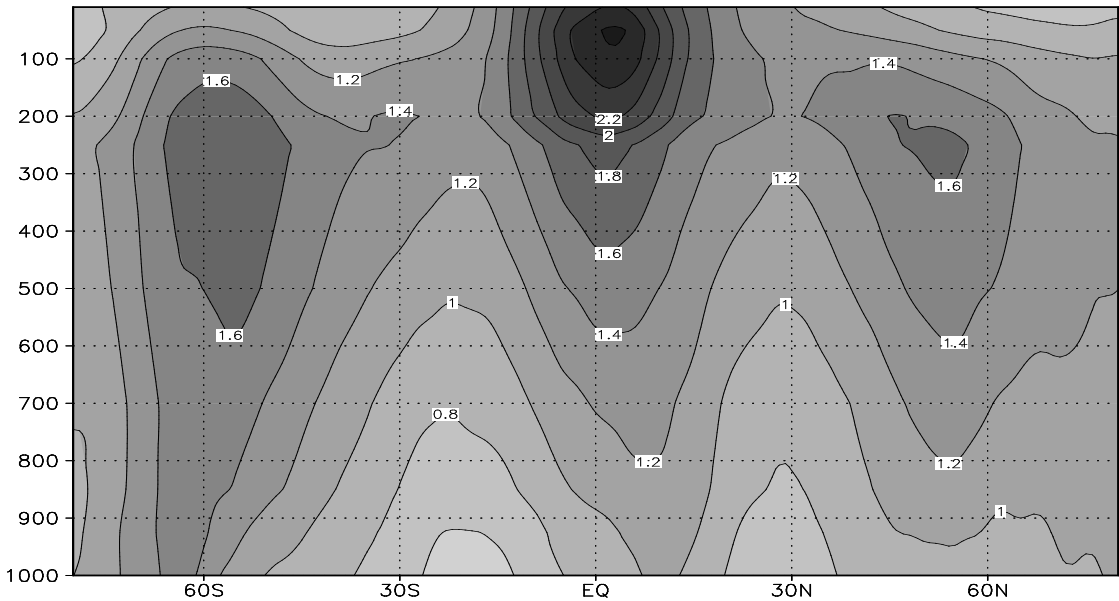
(a) ETR



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(b) ET_3DR



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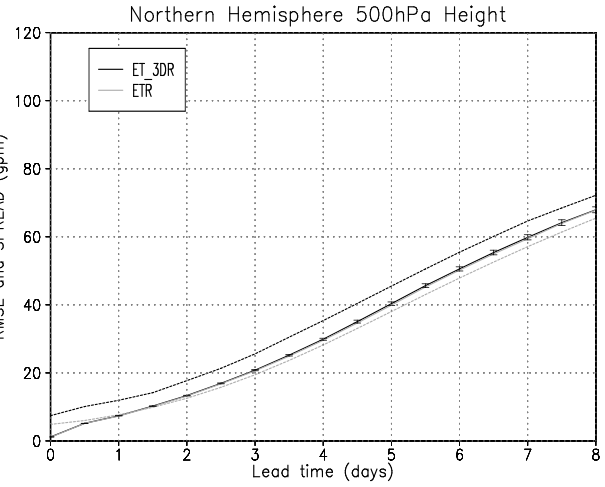
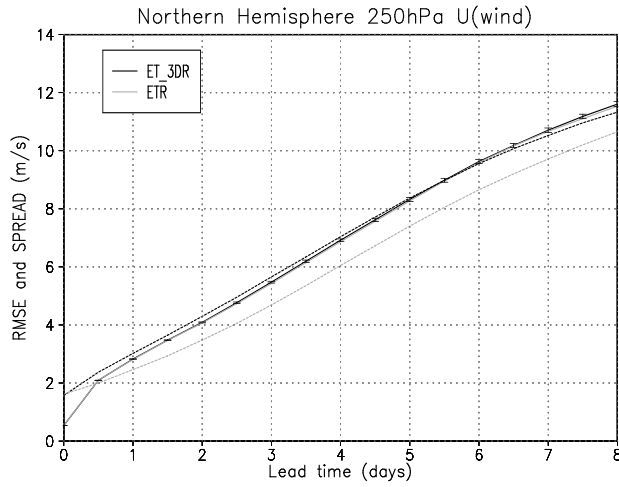
676 Fig. 8 The zonal average of the square root of total energy (m s^{-1}) of initial perturbations for the
677 (a) ETR and (b) ET_3DR experiments at the 500hPa level as a mean for the period 11 September
678 - 30 November 2012.

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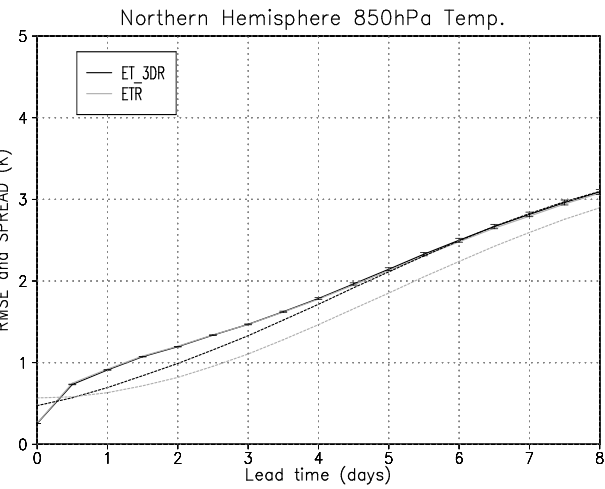
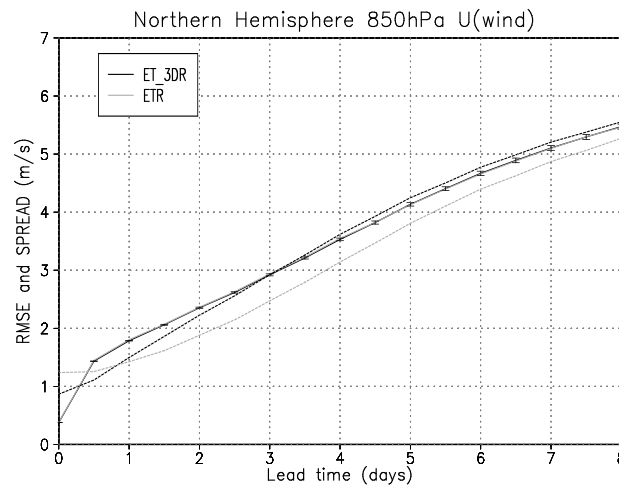


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684 Fig. 9 The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U250, (b) Z500,
 685 (c) U850 and (d) T850 over the NH. The vertical bars represent the 95% confidence interval
 686 from a paried block bootstrap.

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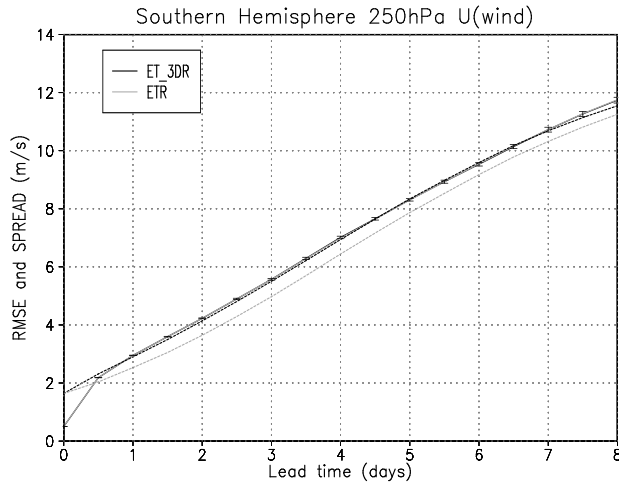
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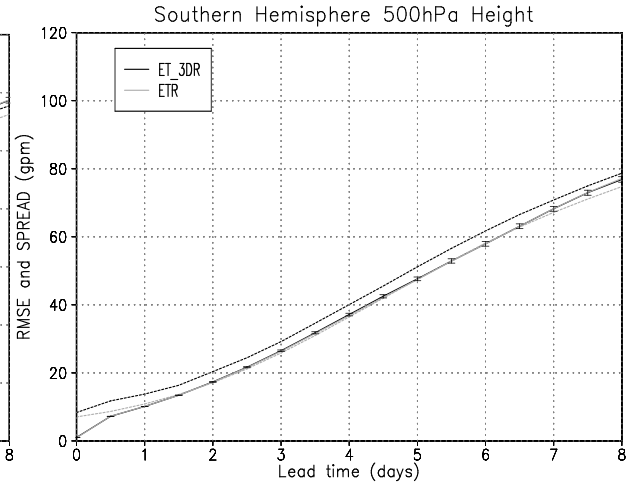
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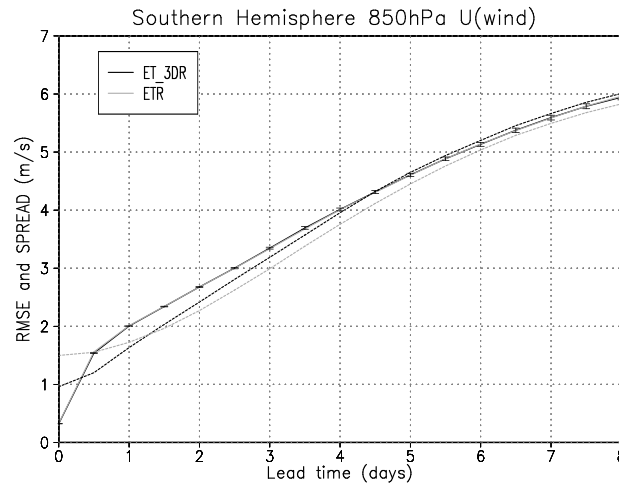
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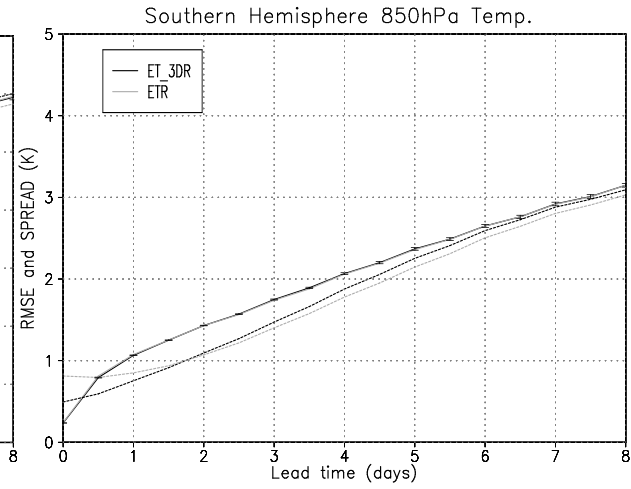
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695 Fig. 10 Same as in Fig. 9, but over the SH.

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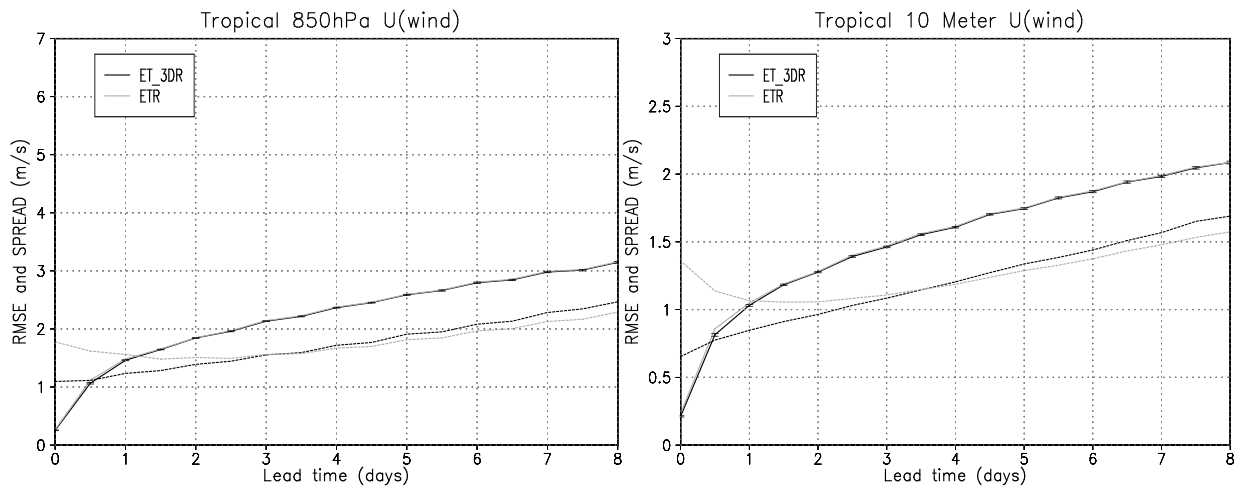
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704 Fig. 11 The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U850 and (b)
705 U10m over the TR for the period 11 September - 30 November 2012. The vertical bars represent
706 the 95% confidence interval from a paired block bootstrap.

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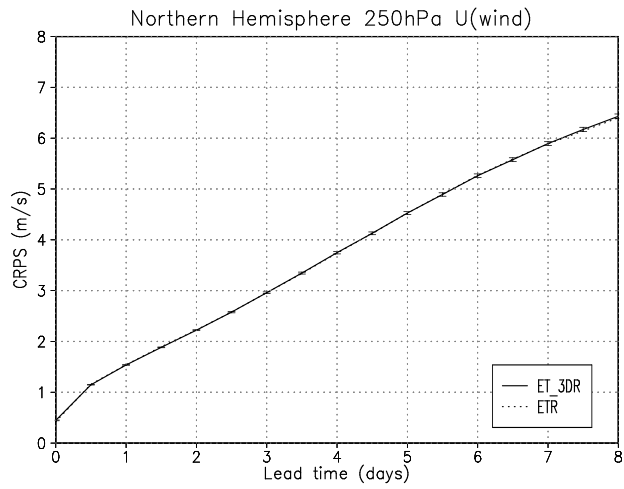
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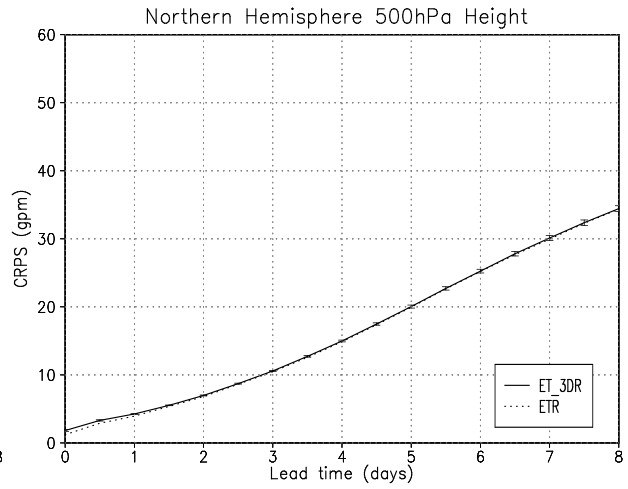
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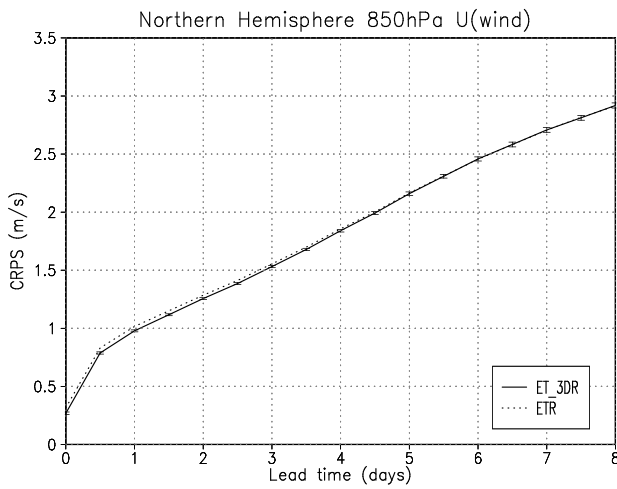
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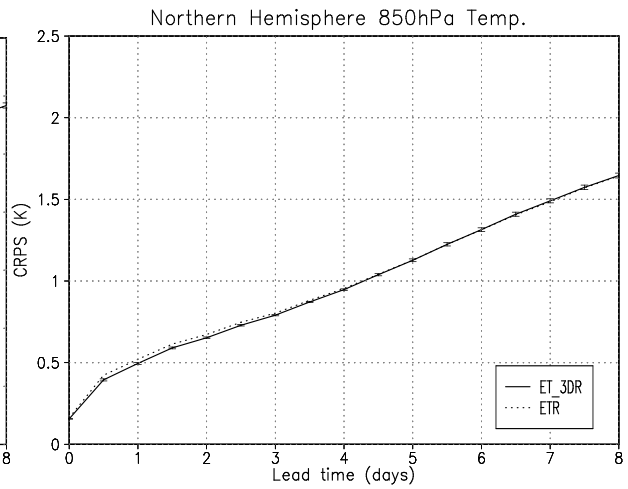
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723 Fig. 12 The CRPS for (a) U250, (b) Z500, (c) U850 and (d) T850 over the NH for the period 11
724 September - 30 November 2012. The vertical bars represent the 95% confidence interval from a
725 paried block bootstrap.

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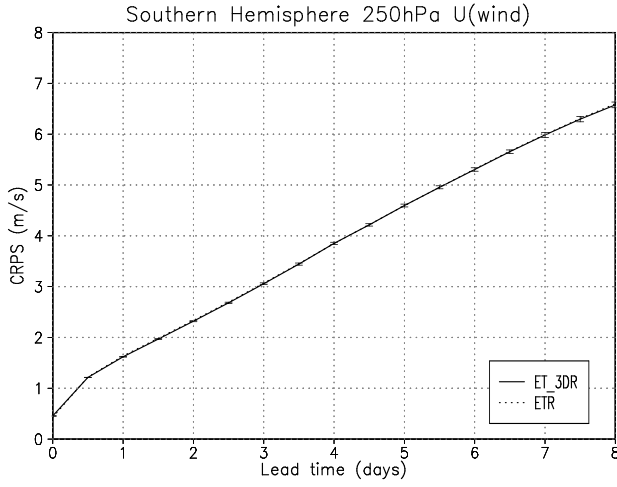
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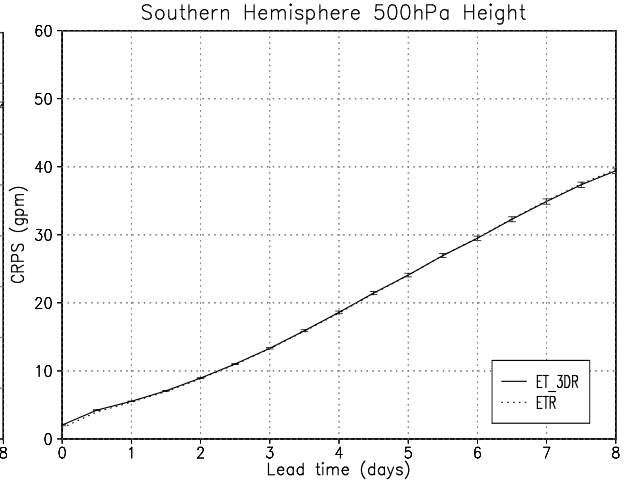
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(a)



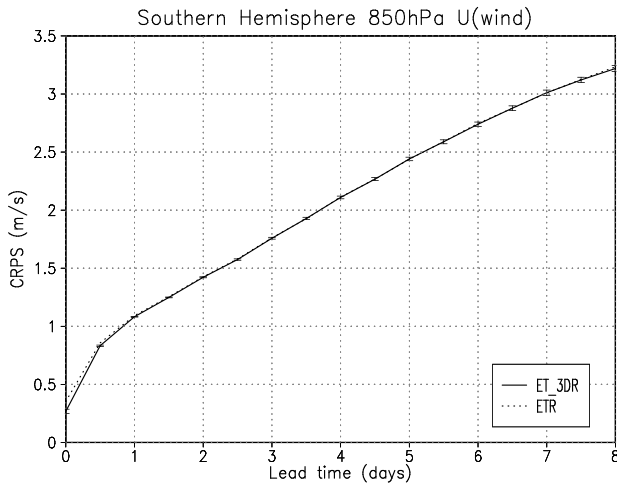
(b)



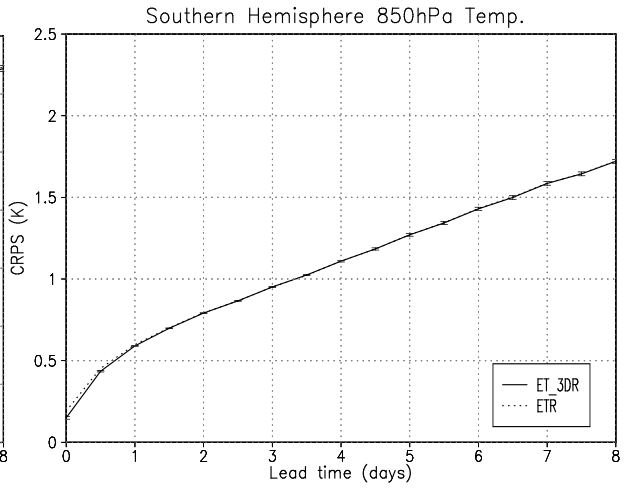
731

732

(c)



(d)



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734 Fig. 13 Same as in Fig. 12, but over the SH.

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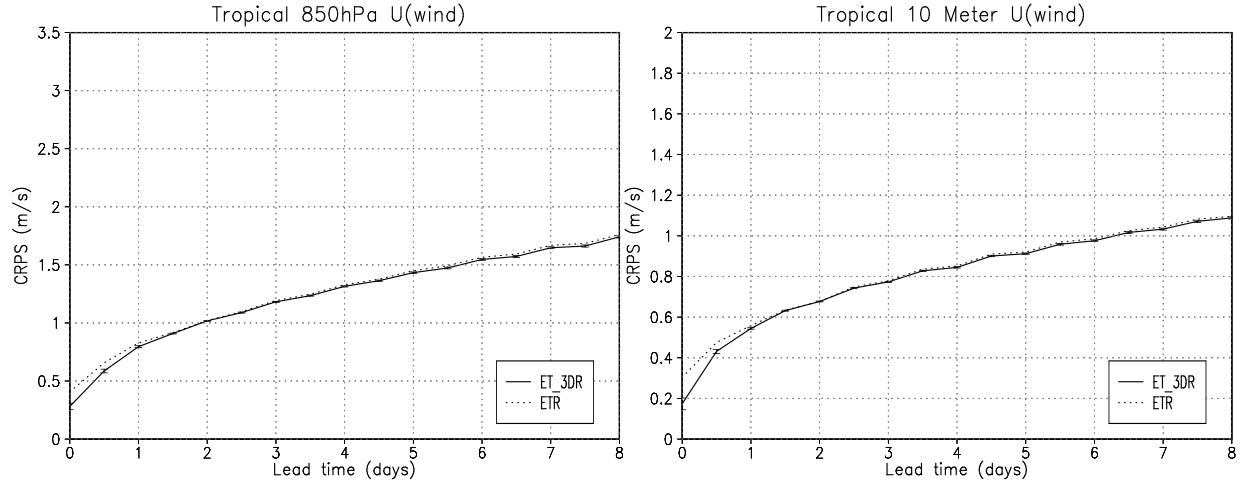
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(a)

(b)



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743 Fig.14 The CRPS for (a) U850 and (b) U10m over the TR for the period 11 September - 30
744 November 2012. The vertical bars represent the 95% confidence interval from a paired block
745 bootstrap.

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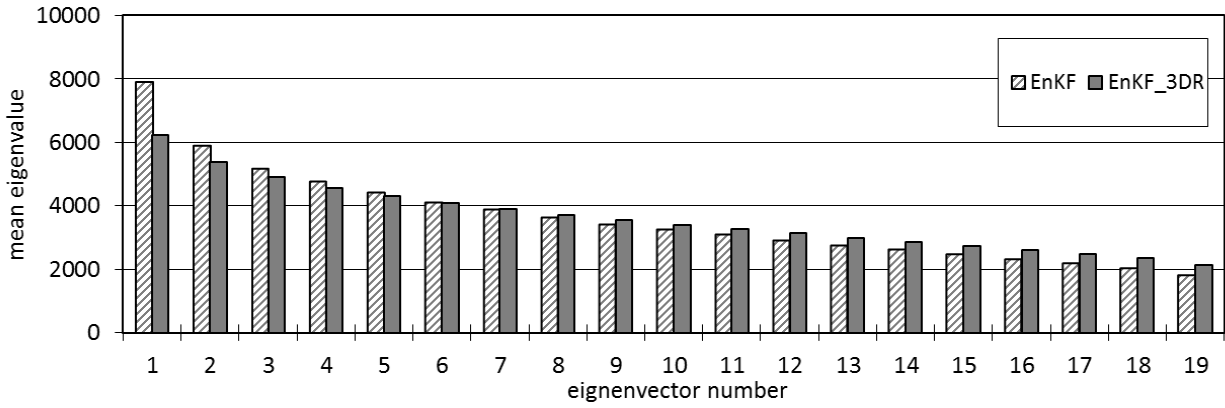
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756 Fig. 15 The mean eigenvalue spectra of the covariance matrix of the initial perturbations for
 757 Z500 during the period 11 September - 30 November 2012 over the globe.

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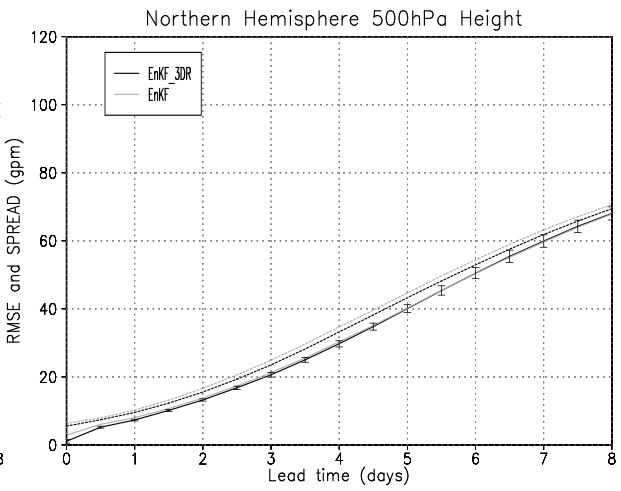
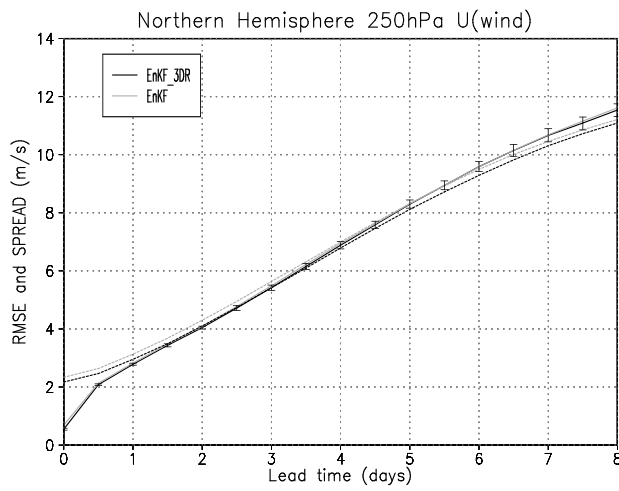
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(a)

(b)

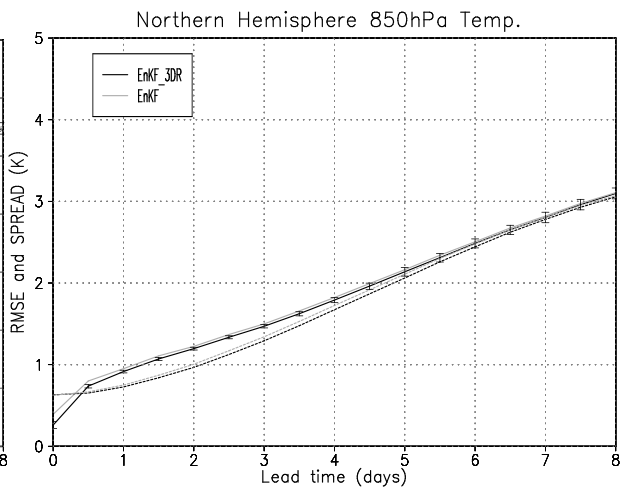
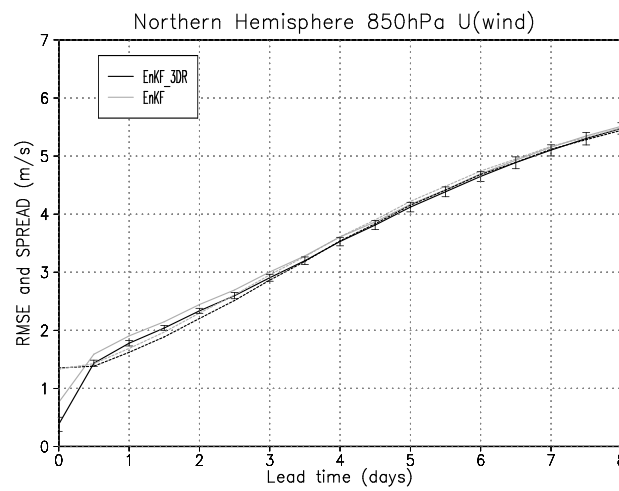


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(c)

(d)



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777 Fig. 16 The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U250, (b) Z500,
 778 (c) U850, and (d) T850 over the NH. The vertical bars represent the 95% confidence interval
 779 from a paried block bootstrap.

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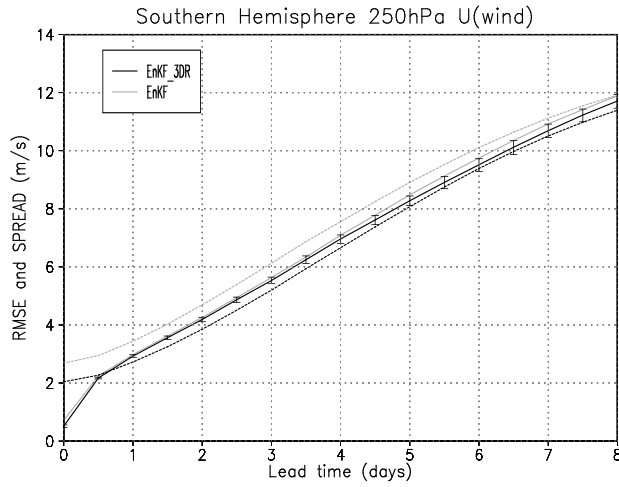
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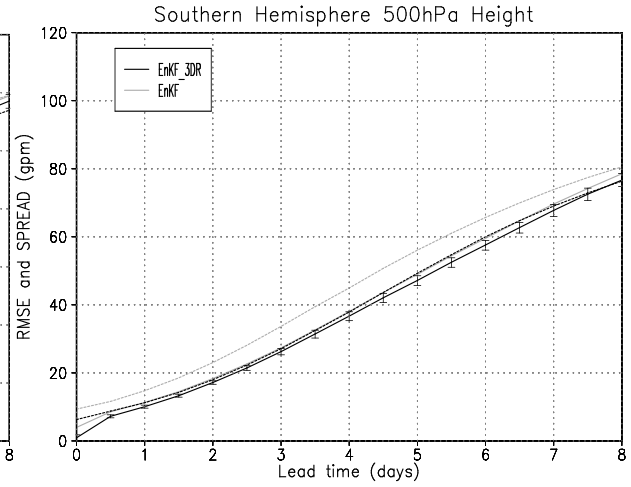
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(a)



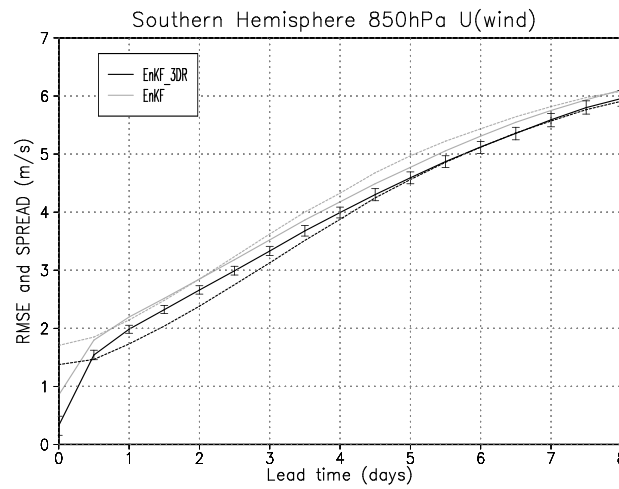
(b)



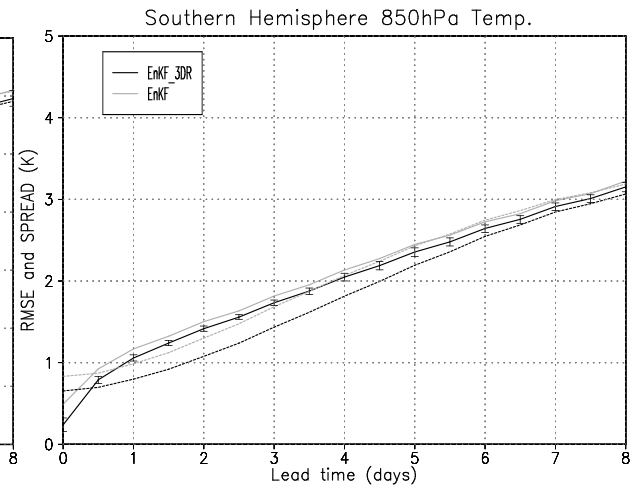
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(c)



(d)



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788 Fig. 17 Same as in Fig. 16, but over the SH.

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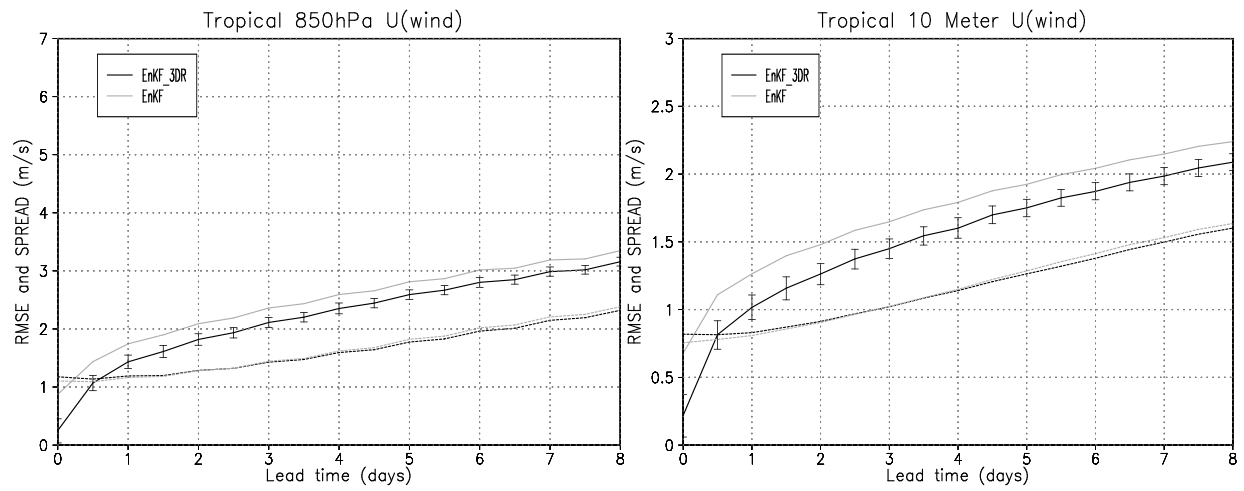
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(a)

(b)



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797 Fig. 18 The ensemble mean RMSE (solid) and ensemble spread (dashed) for (a) U850 and (b)
798 U10m over the TR for the period 11 September - 30 November 2012. The vertical bars represent
799 the 95% confidence interval from a paired block bootstrap.

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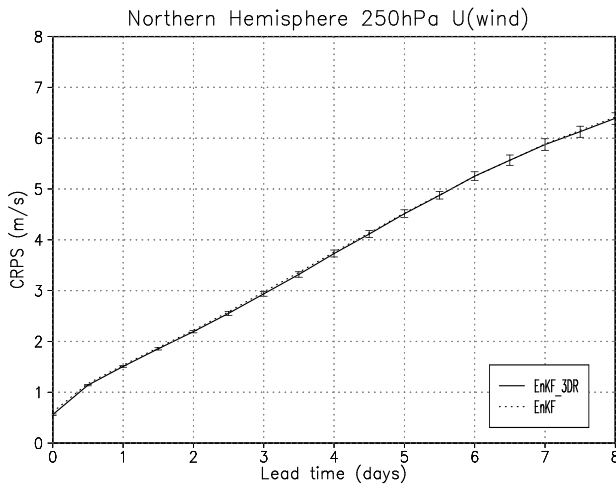
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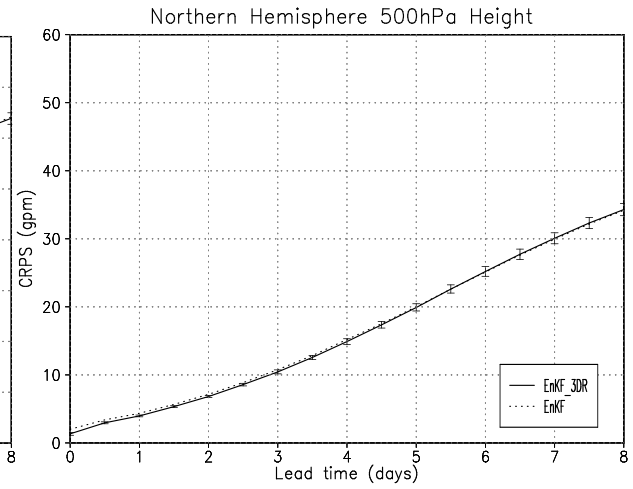
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(a)



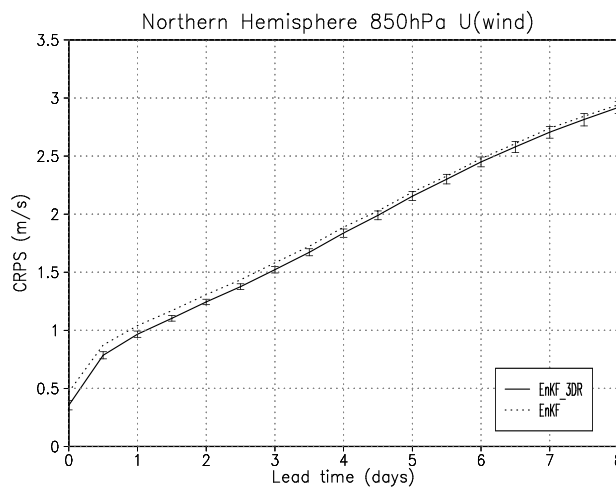
(b)



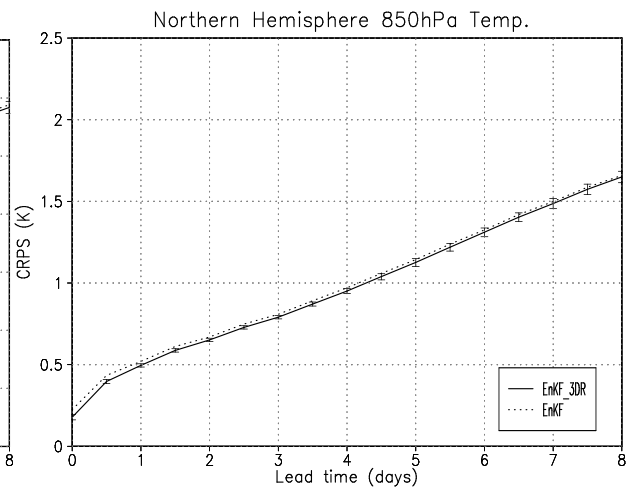
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(c)



(d)



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816 Fig. 19 The CRPS for (a) U250, (b) Z500, (c) U850 and (d) T850 over the NH for the period 11

817 September - 30 November 2012. The vertical bars represent the 95% confidence interval from a

818 paired block bootstrap.

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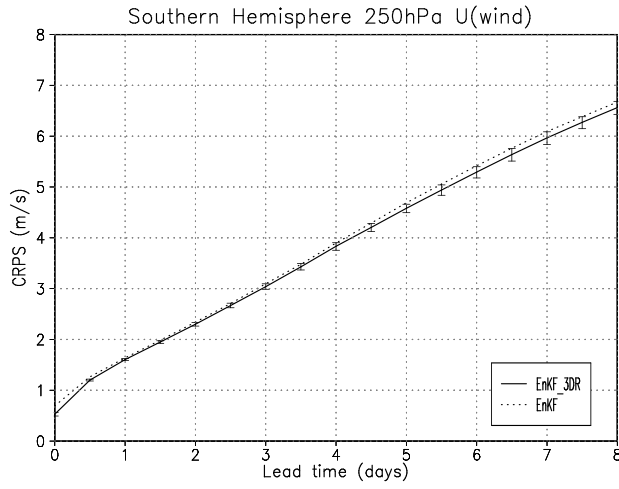
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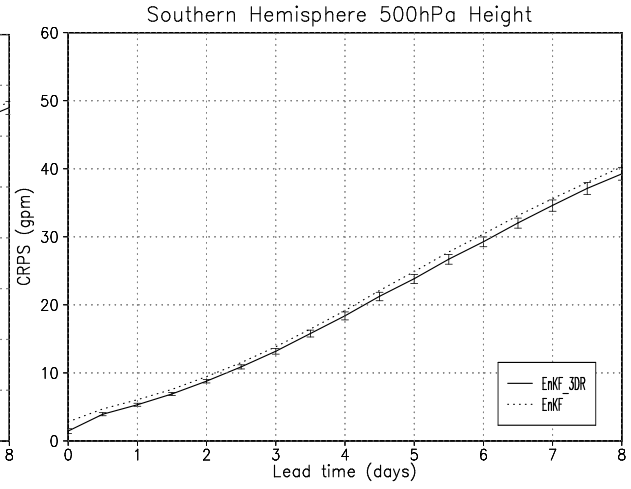
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(a)



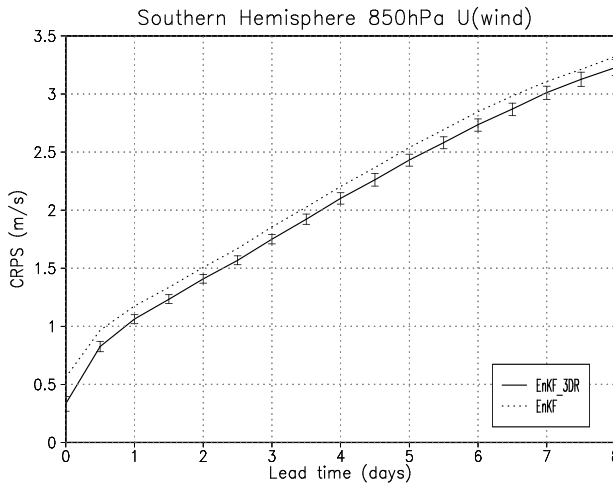
(b)



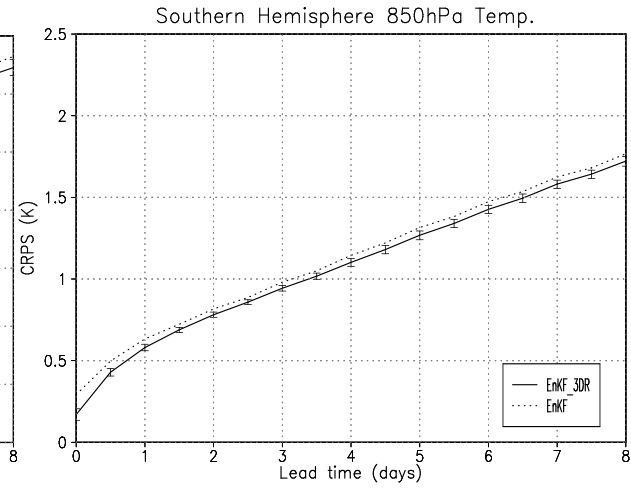
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(c)



(d)



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827 Fig. 20 Same as in Fig. 19, but over the SH.

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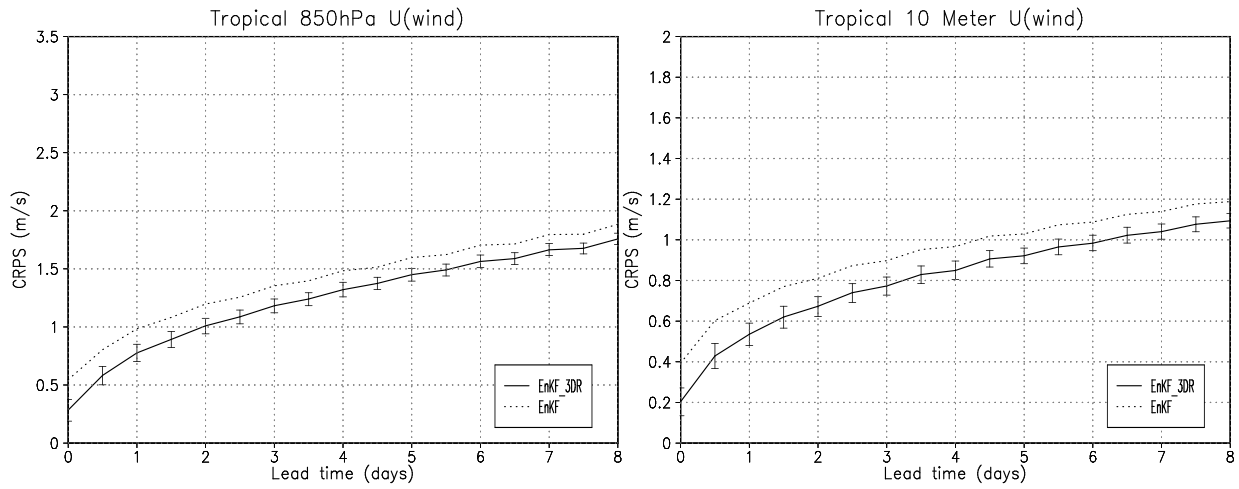
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(a)

(b)



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836 Fig. 21 The CRPS for (a) U850 and (b) U10m over the TR for the period 11 September - 30
837 November 2012. The vertical bars represent the 95% confidence interval from a paired block
838 bootstrap.

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